



A Dust Aggregate Model

Based on

Numerical Simulations of Aggregate Collisions

ダストアグリゲイト同士の
DEM(?)衝突シミュレーション

和田 浩二

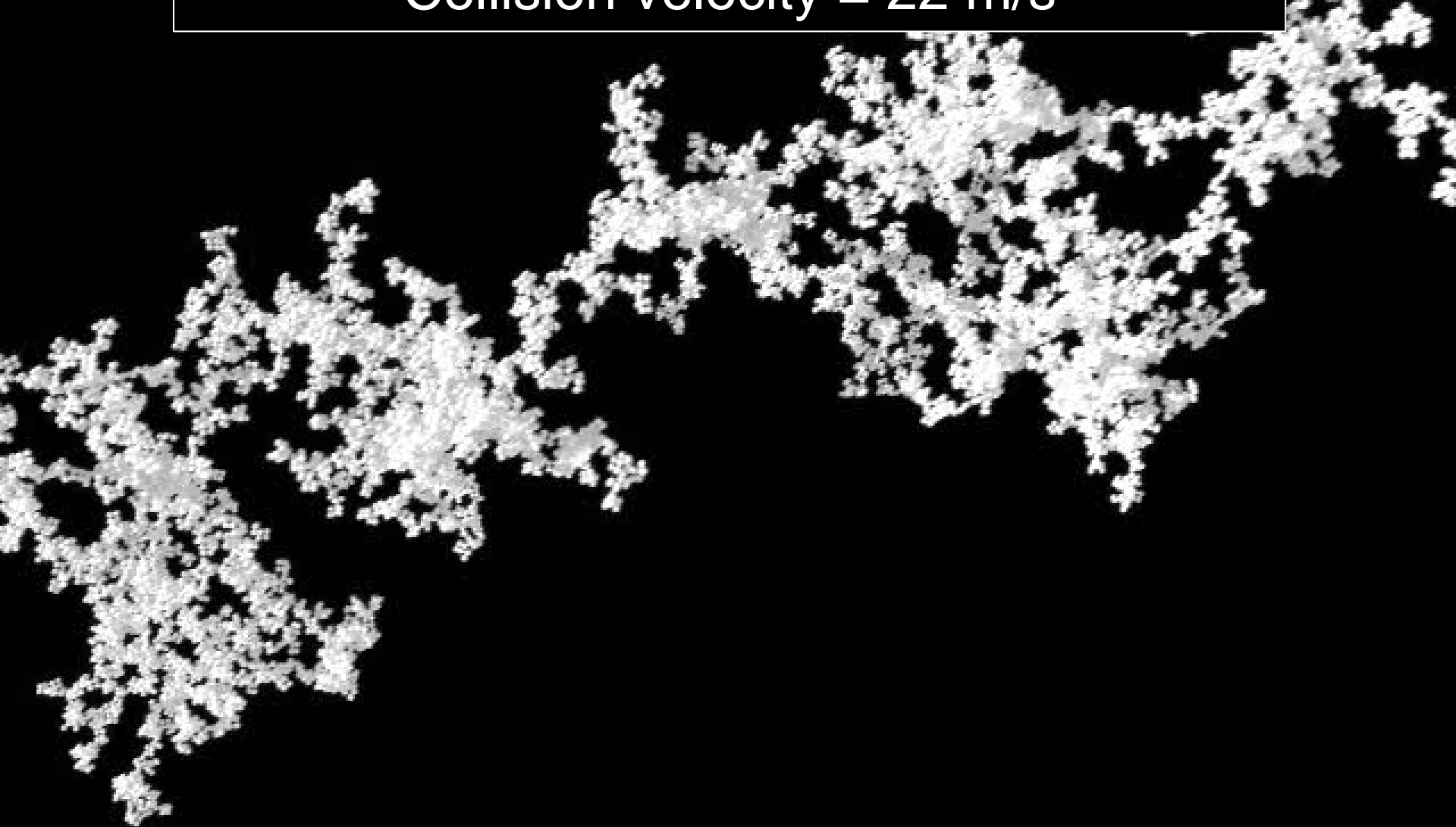
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A collision of BCCAs
8192+8192 ice particles ($r=0.1\mu\text{m}$, $\xi_c = 8\text{\AA}$)
Collision velocity = 22 m/s



Background



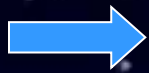
Collisional growth of dust
($< \mu\text{m}$)



Planetesimal formation
($> \text{km}$)

Structure evolution of dust aggregates in protoplanetary disks

When and how are aggregates compressed and/or disrupted ?



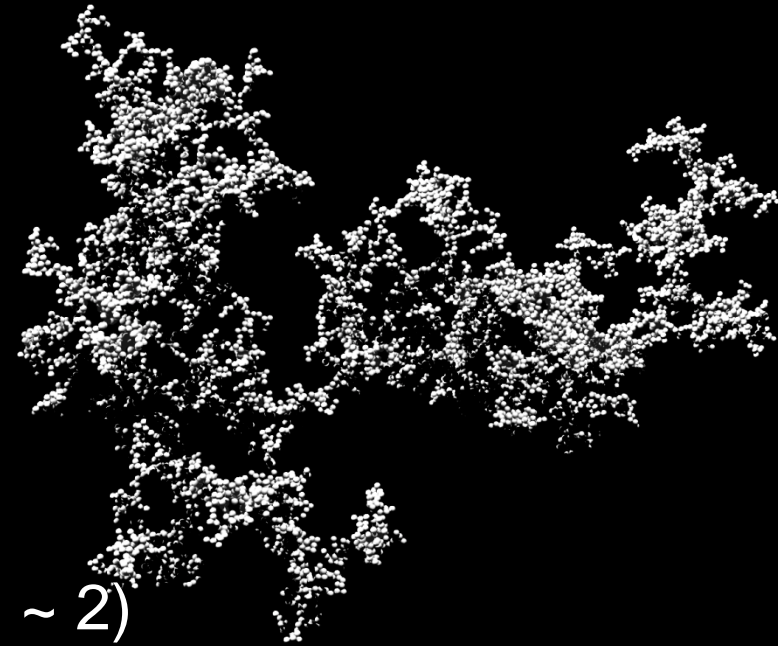
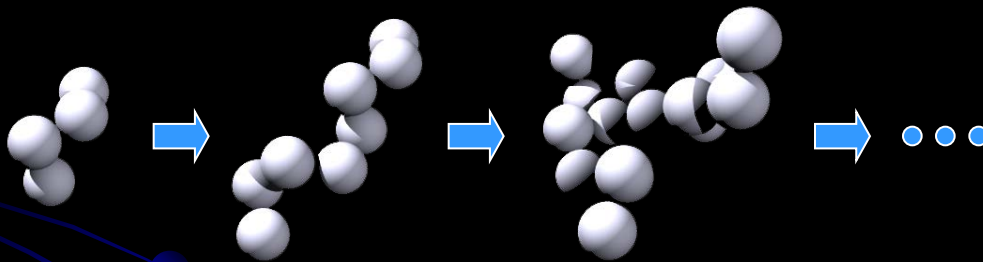
Numerical simulation of dust aggregate collisions!

Ballistic Cluster-Cluster Aggregation (BCCA)



✓ In the early growth stage, **undeformed BCCAs** are formed because of their low collision velocity ($< \text{mm/s}$)

- A series of hit-and sticks of comparable aggregates



- **Fluffy** structure (fractal dimension $< \sim 2$)

How are the BCCA structures compressed?

Dominik & Tielens 1997;

Wada et al. 2007, 2008; Suyama et al. 2008

Background



Collision velocity of dust
in protoplanetary disks $<$ several 10 m/s

e.g., $<$ ~ 50 m/s (Hayashi model, without turbulence)



Is it possible for dust to grow through collisions?
To what extent is dust compressed?

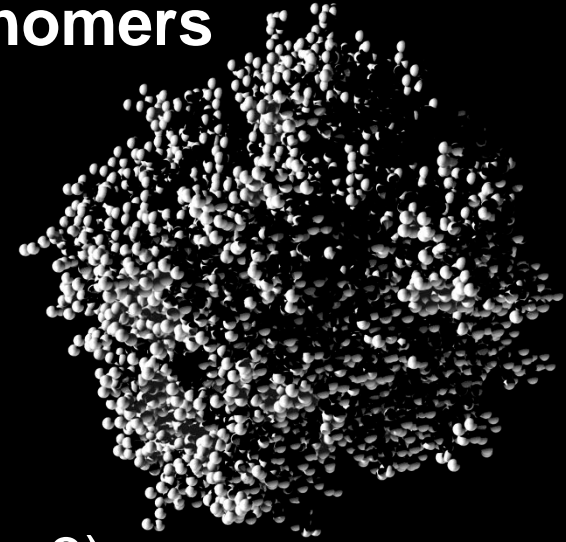
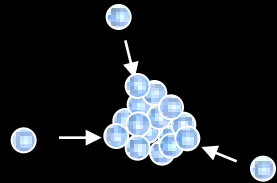
Experimental: Blum & Wurm 2000, Wurm et al. 2005

Numerical: Dominik & Tielens 1997, Wada et al. 2008; 2009

Ballistic Particle-Cluster Aggregation (BPCA)



- Formed by one-by-one sticking of monomers



- **Compact** structure (fractal dimension ~ 3)

Dust should be compact

in high velocity collisions causing their disruption

Collisions of BPCA clusters

→ implication for growth and disruption of compact dust

Objective



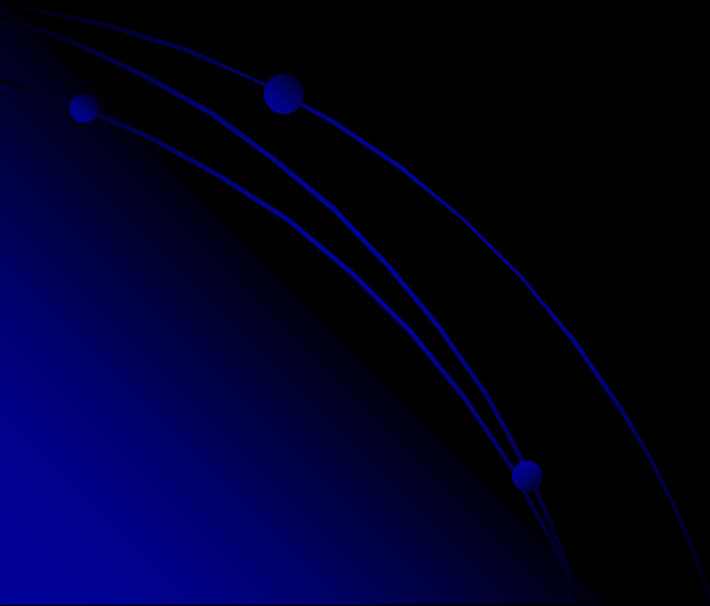
To construct a structural evolution model of dust aggregates by numerical simulations of aggregate collisions

Collisions of BCCA & BPCA clusters

- Compression process (BCCAs)
 - Gyration radius → Degree of compression
- High-velocity collisions (BPCAs)
 - Number of particles in the largest remnant → Growth efficiency
 - Coordination numbers in the largest remnant → Degree of compression



Simulation Method



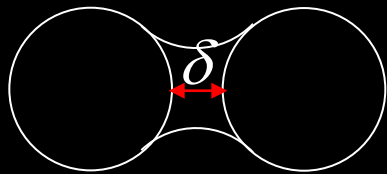
Grain interaction model

Johnson, Kendall and Roberts (1971)
 Johnson (1987), Chokshi et al. (1993)
 Dominik and Tielens (1995,96)
 Wada et al. (2007)

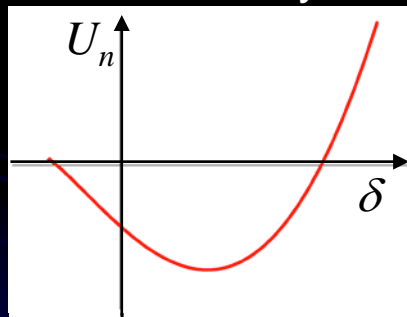


Elastic spheres having surface energy

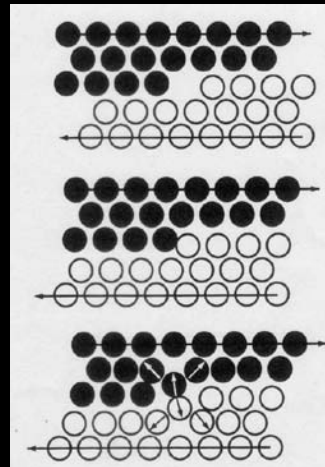
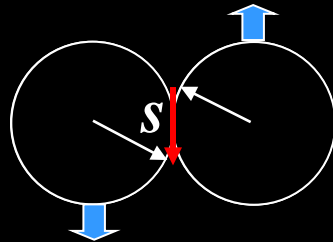
Normal



JKR theory

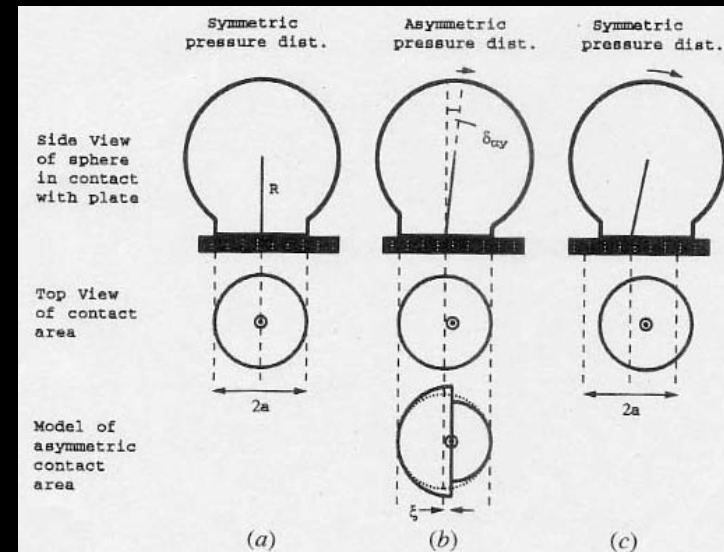
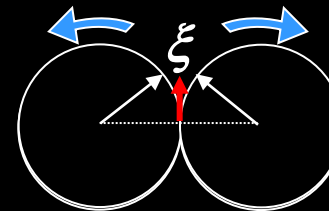


Sliding



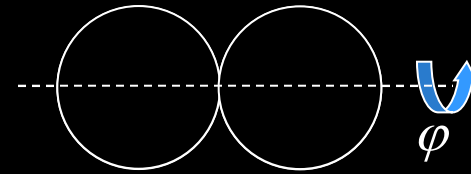
(Dominik & Tielens 1996)

rolling



(Dominik & Tielens 1995)

twisting



Critical sticking velocity:
 exp. $\sim 10 \times$ theo.!?

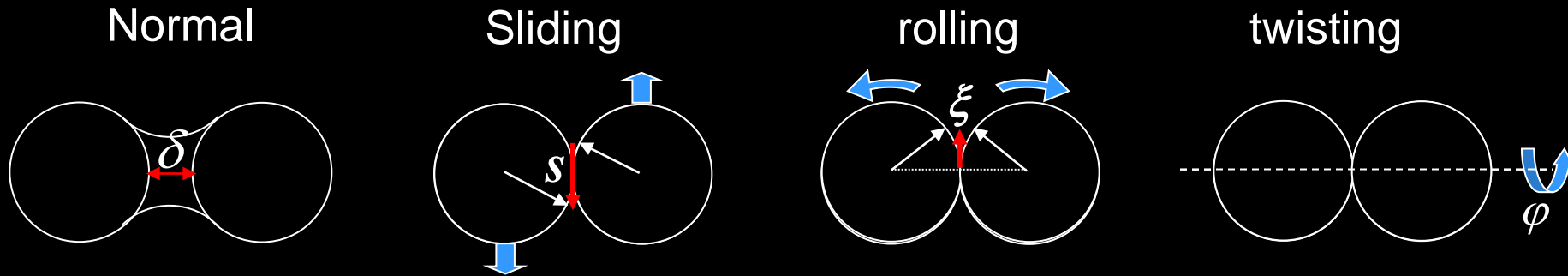
JKR and rolling resistance have been tested with experiments using $\sim 1 \mu\text{m}$ SiO_2 particles. (Heim et al. 1999; Poppe et al. 2000; Blum & Wurm 2000)

Grain interaction model

Johnson, Kendall and Roberts (1971)
 Johnson (1987), Chokshi et al. (1993)
 Dominik and Tielens (1995,96)
 Wada et al. (2007)



Elastic spheres having surface energy



Contact & Separation

$s, \xi, \varphi >$ critical displacements

→ Energy dissipation

- Critical slide $s_{crit} \sim 1.5 \text{ \AA}$ (for 0.2 \mu m quartz)
- Critical roll $\xi_{crit} \sim 2 \text{ \AA}$ (or $\sim 30 \text{ \AA}$ (Heim et al., 1999))
- Critical twist $\varphi_{crit} \sim 1^\circ$

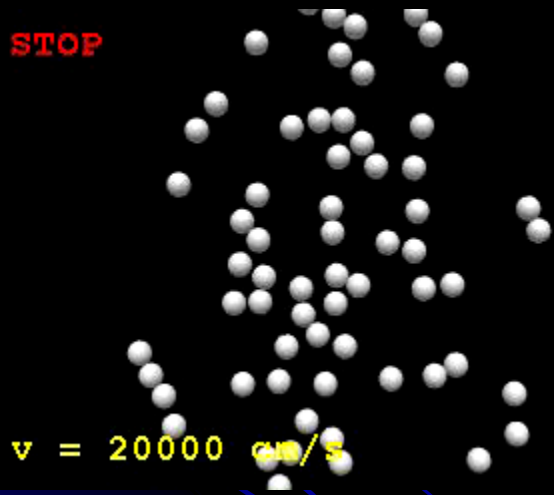
E_{break} : Energy to break a contact

E_{roll} : Energy to roll a pair of grains by 90°

A classical study

Dominik and Tielens (1997)

Each grain motion is directly calculated, taking into account particle interactions



✓ modeling grain interactions seriously

Limitations:
D&T "recipe"

- 2-D, Head-on collision
- $E_{\text{impact}} = \begin{cases} \sim n_k E_{\text{roll}} & \rightarrow \text{Max. compression} \\ > 10 n_k E_{\text{break}} & \rightarrow \text{Catastrophic disruption} \end{cases}$
- Small size (40+40 grains)
- Initial structure: only 1 type

E_{roll} : Energy to roll a grain by 90°

E_{break} : Energy to break a contact

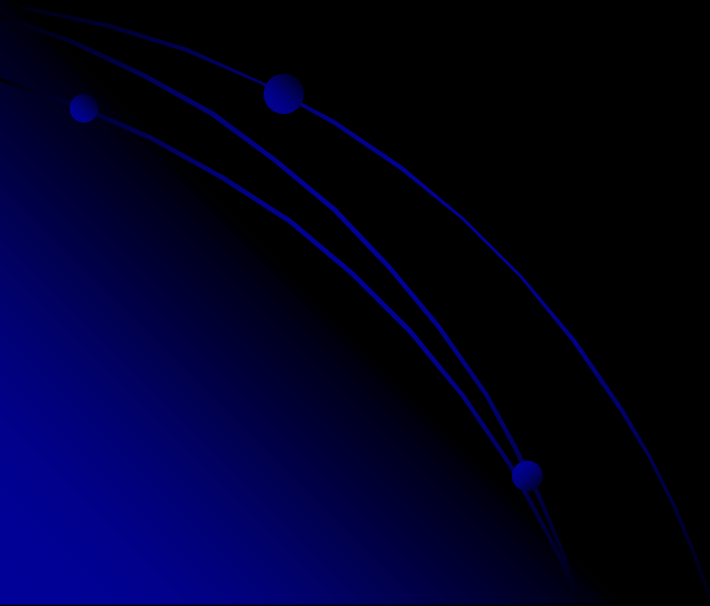
n_k : Number of contacts in initial aggregates

Confirmed by experiments
(Blum & Wurm 2000)



Collisions between BCCA clusters

: Compression process



Initial Conditions and Parameters

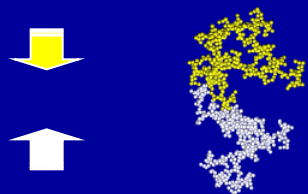


Collisions of BCCA clusters

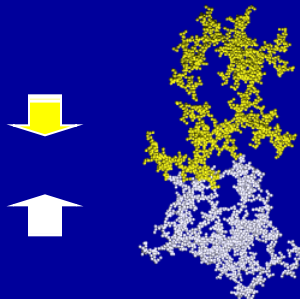
✓ BCCA clusters are

- composed of **512, 2048, or 8192** particles (10 types randomly produced)
- impacted by **head-on** collision

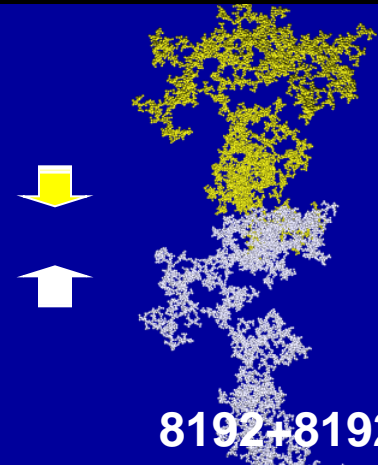
Results are averaged



512+512



2048+2048



8192+8192

✓ particle : radius = $0.1 \mu\text{m}$,

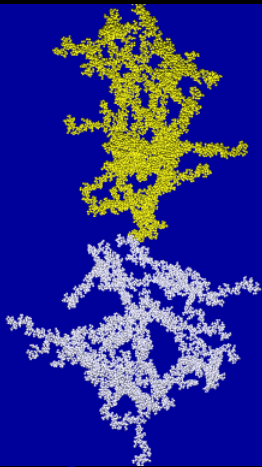
Ice ($E = 7 \text{ GPa}$, $\nu = 0.25$, $\gamma = 100 \text{ mJ/m}^2$)

SiO₂ ($E = 54 \text{ GPa}$, $\nu = 0.17$, $\gamma = 25 \text{ mJ/m}^2$)

✓ Critical rolling displacement : $\xi_{\text{crit}} = 2, 8, 30 \text{ \AA}$

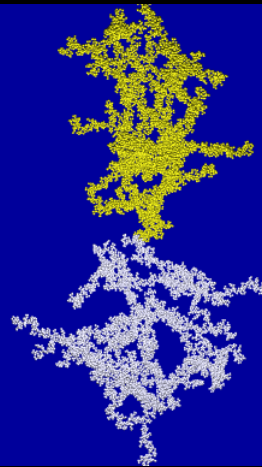
Example of simulations

Ice, 8192 + 8192, $\xi_{\text{crit}} = 8 \text{ \AA}$



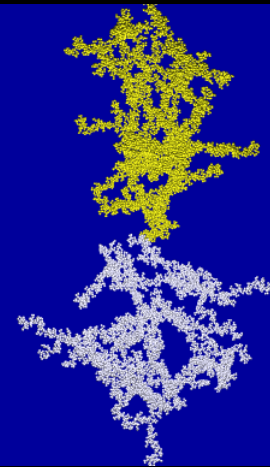
$$E_{\text{impact}} \sim 0.7 E_{\text{roll}}$$

$$V_{\text{impact}} = 0.2 \text{ m/s}$$



$$E_{\text{impact}} \sim 0.3 n_k E_{\text{roll}}$$

$$V_{\text{impact}} = 17 \text{ m/s}$$

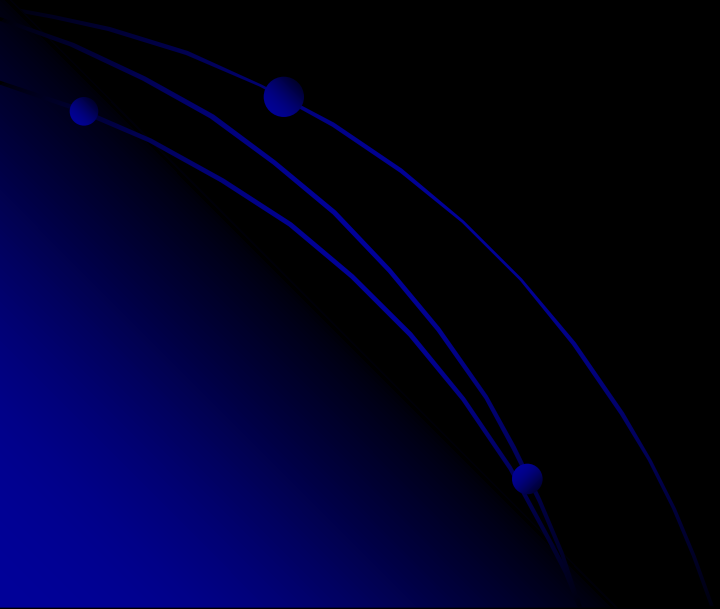


$$E_{\text{impact}} \sim 13 n_k E_{\text{break}}$$

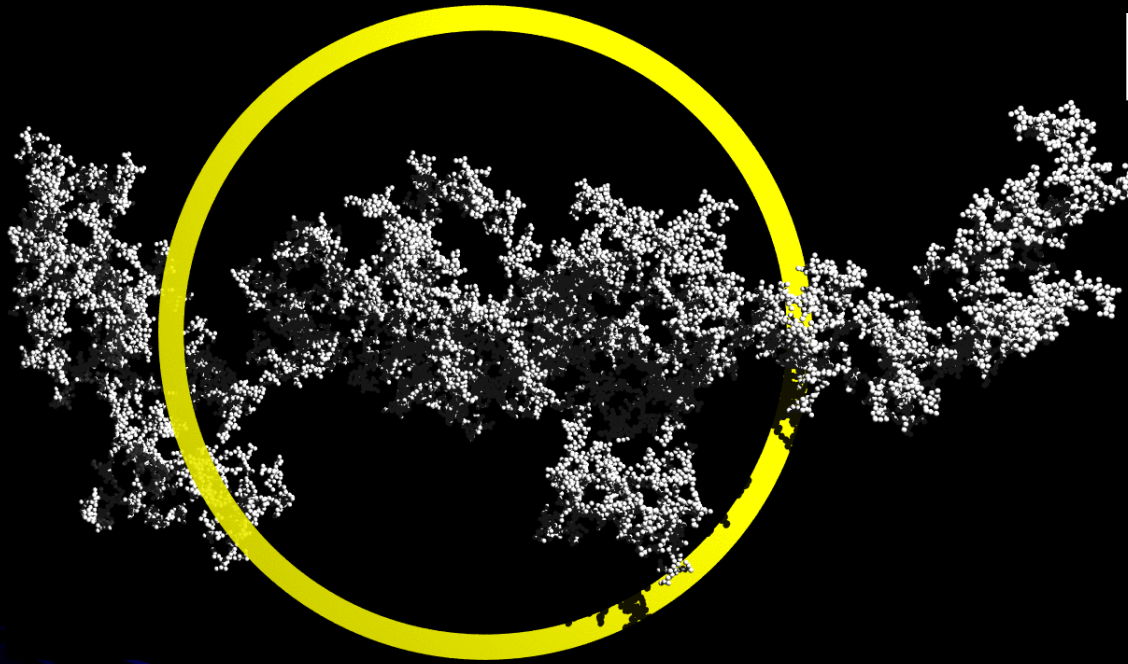
$$V_{\text{impact}} = 39 \text{ m/s}$$



Numerical Results on Gyration Radius



Gyration radius r_g : *compression process*



$$E_{\text{impact}} \sim 0.01 E_{\text{roll}}$$

Impact velocity: 0.024 m/s

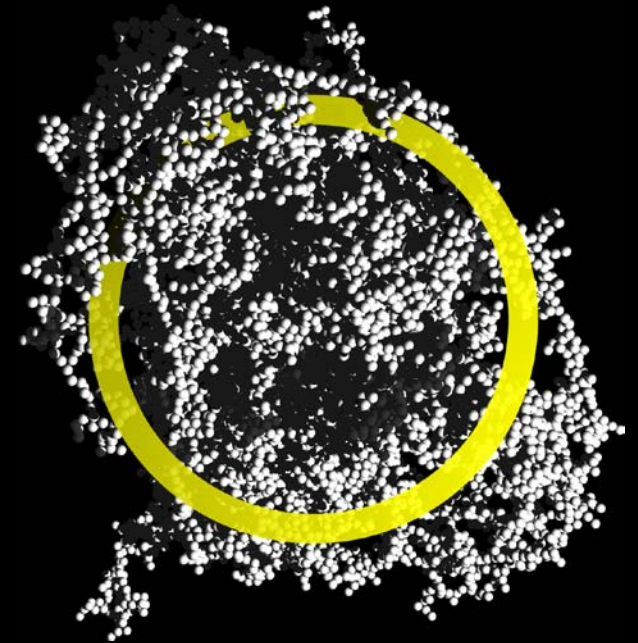
$$E_{\text{impact}} \sim 0.19 N E_{\text{roll}}$$

Impact velocity: 13 m/s

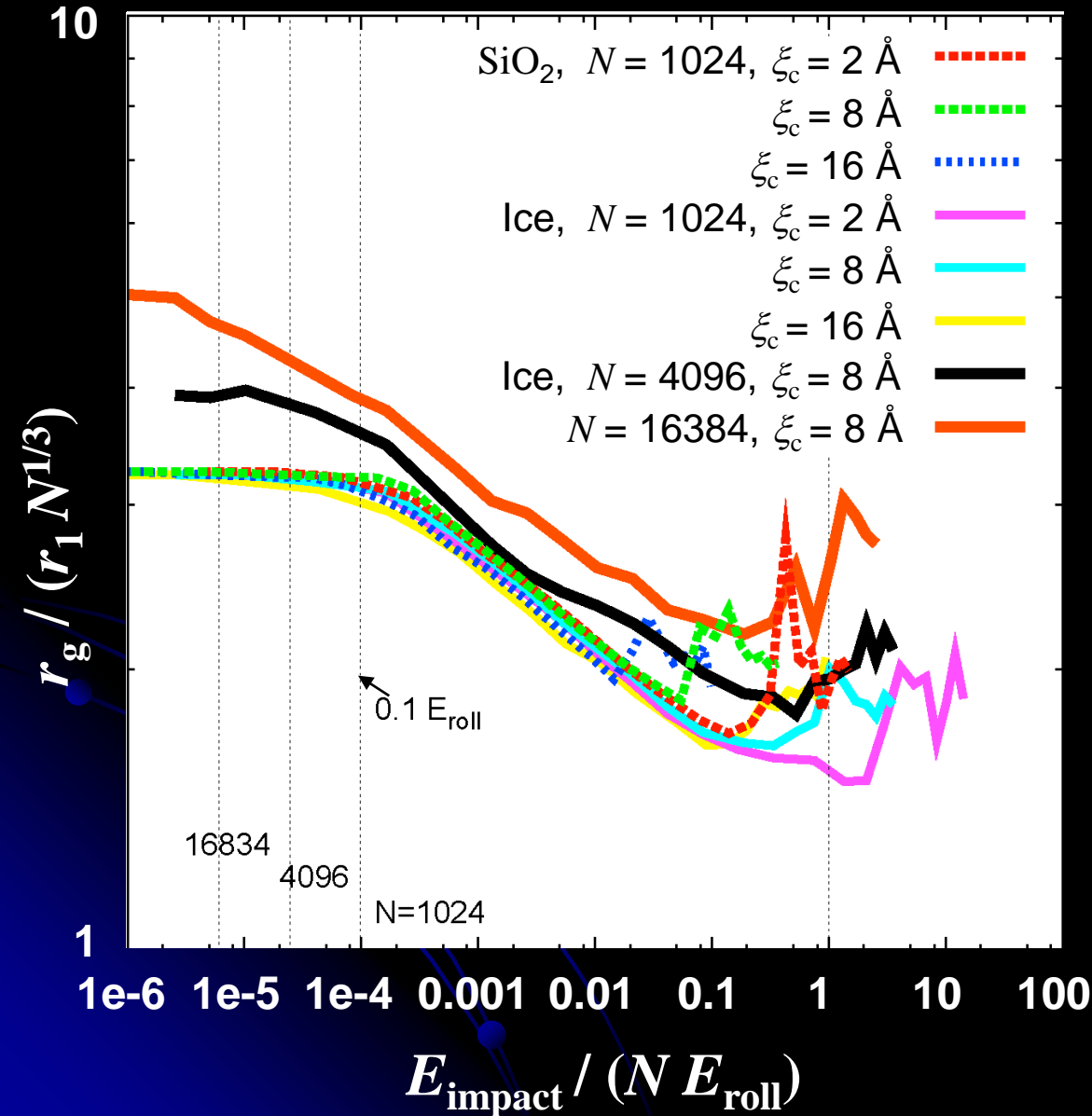
Ice, 8192 + 8192, $\xi_{\text{crit}} = 8 \text{ \AA}$

$$r_g = \sqrt{\frac{1}{N} \sum_i |x_i - x_g|^2}$$

x_g : center of mass



Gyration radius r_g : *compression process*



$\checkmark E_{\text{impact}} \sim (0.1 - 1) N E_{\text{roll}}$

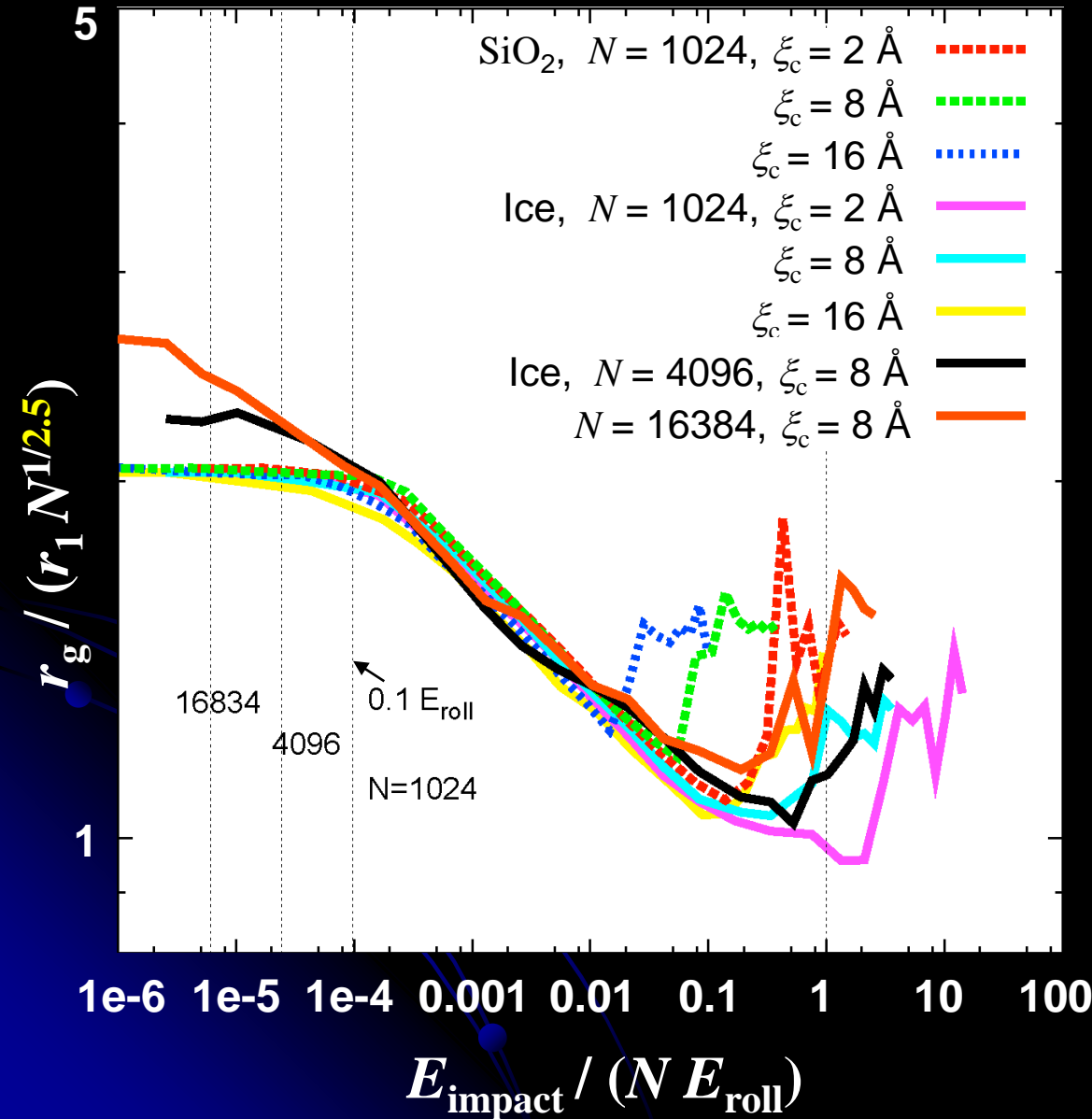
\downarrow

Max. compression

\updownarrow

Consistent with Dominik & Tielens (1997)

Gyration radius r_g : compression process



✓ Scaled by

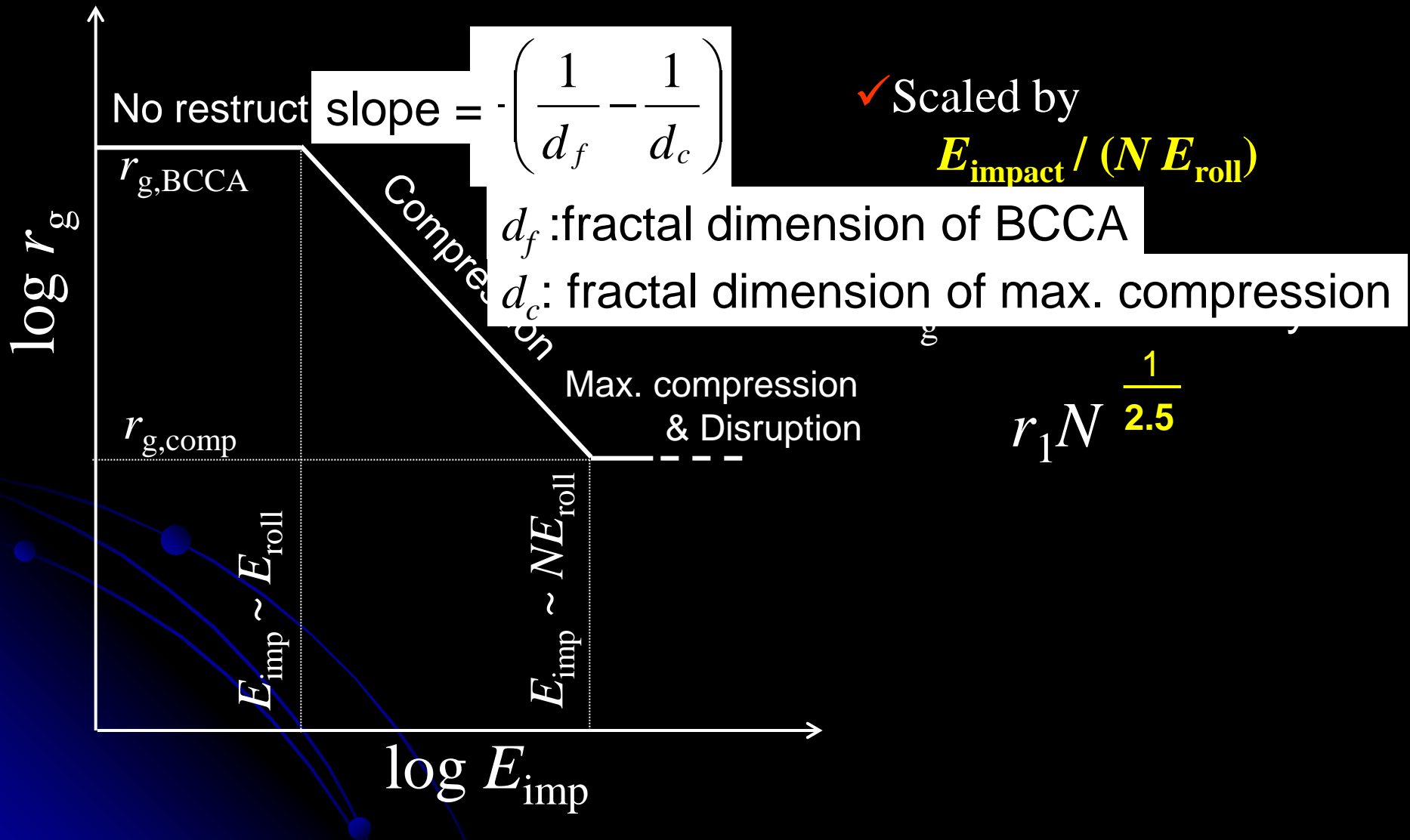
$$E_{\text{impact}} / (N E_{\text{roll}})$$

✓ r_g is normalized by

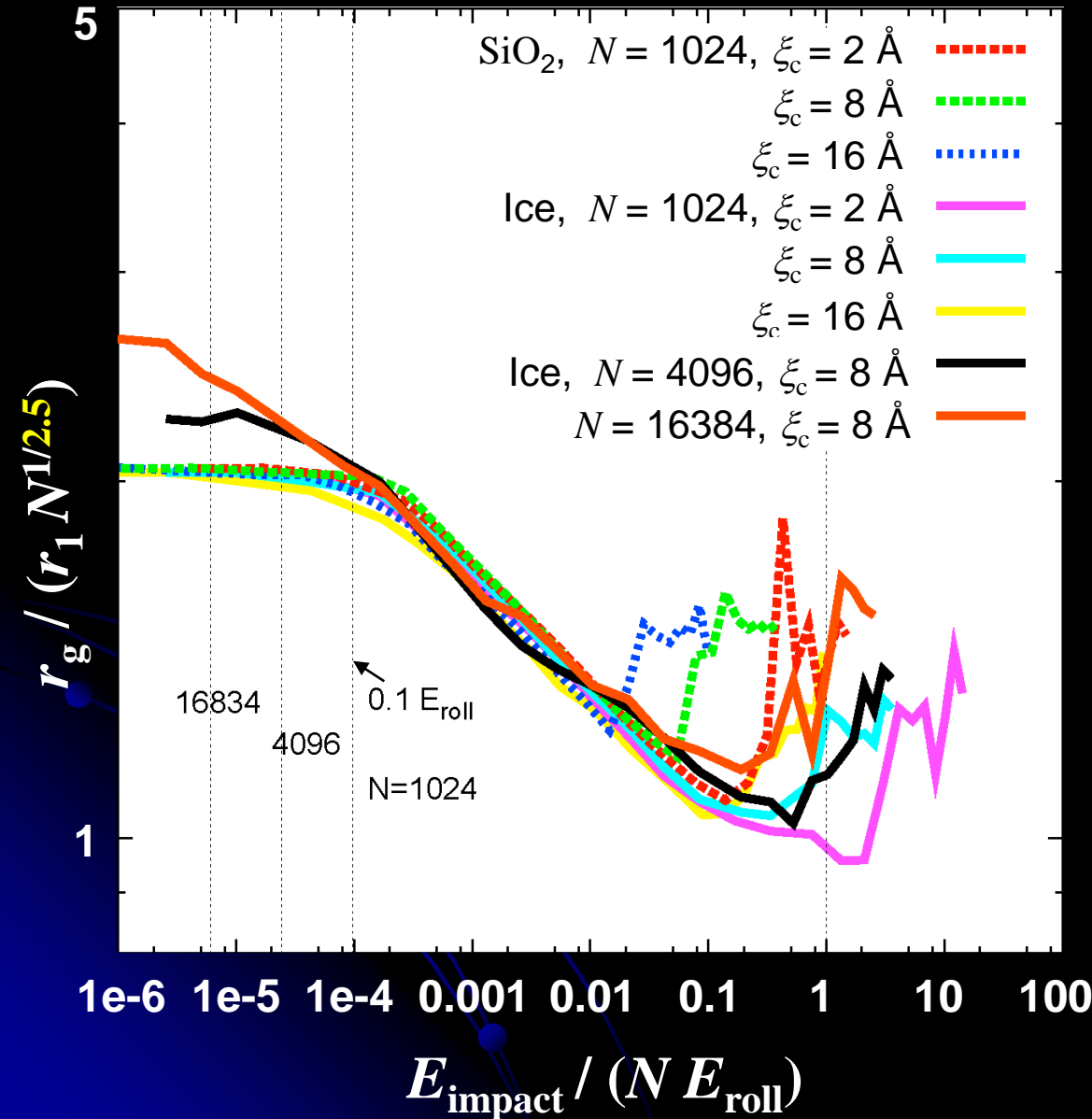
$$r_1 N^{\frac{1}{2.5}}$$



Gyration radius r_g : compression process



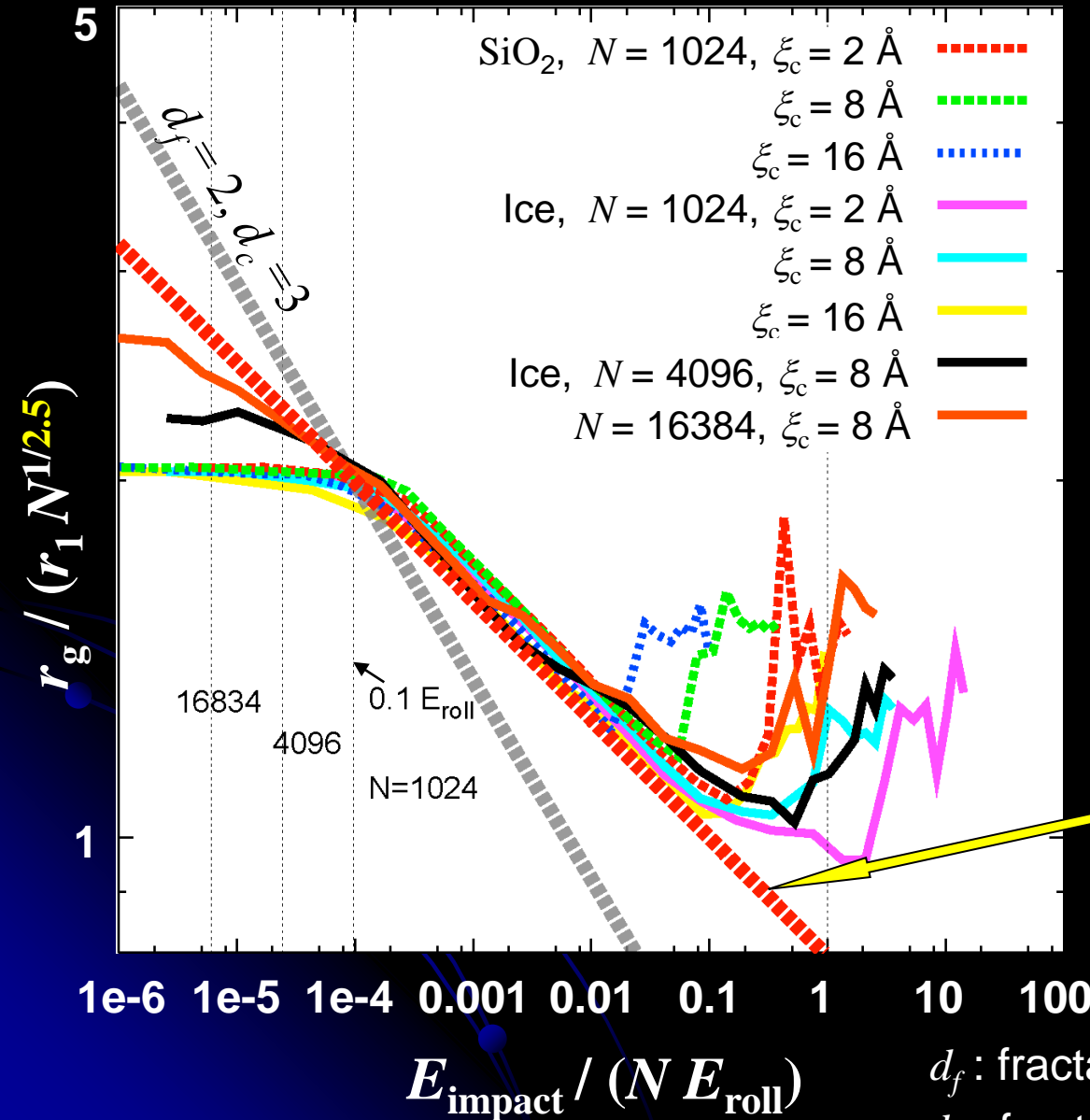
Gyration radius r_g : compression process



✓ Scaled by $E_{\text{impact}} / (N E_{\text{roll}})$

✓ r_g is normalized by $r_1 N^{\frac{1}{2.5}}$

Gyration radius r_g : compression process



✓ Scaled by

$$E_{\text{impact}} / (N E_{\text{roll}})$$

✓ r_g is normalized by

$$r_1 N^{\frac{1}{2.5}}$$

✓ Not fully compressed

$$\frac{r_g}{r_1 N^{1/2.5}} \approx 0.8 \left(\frac{E_{\text{impact}}}{N E_{\text{roll}}} \right)^{-0.1}$$

$$(d_f = 2, d_c = 2.5)$$

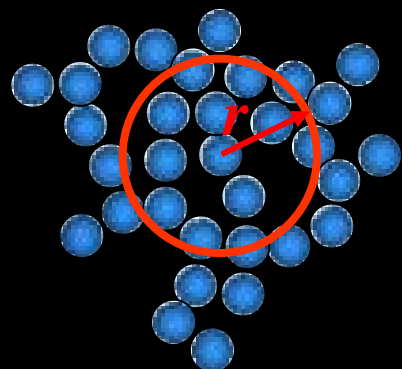
d_f : fractal dimension of BCCA

d_c : fractal dimension of max. compression

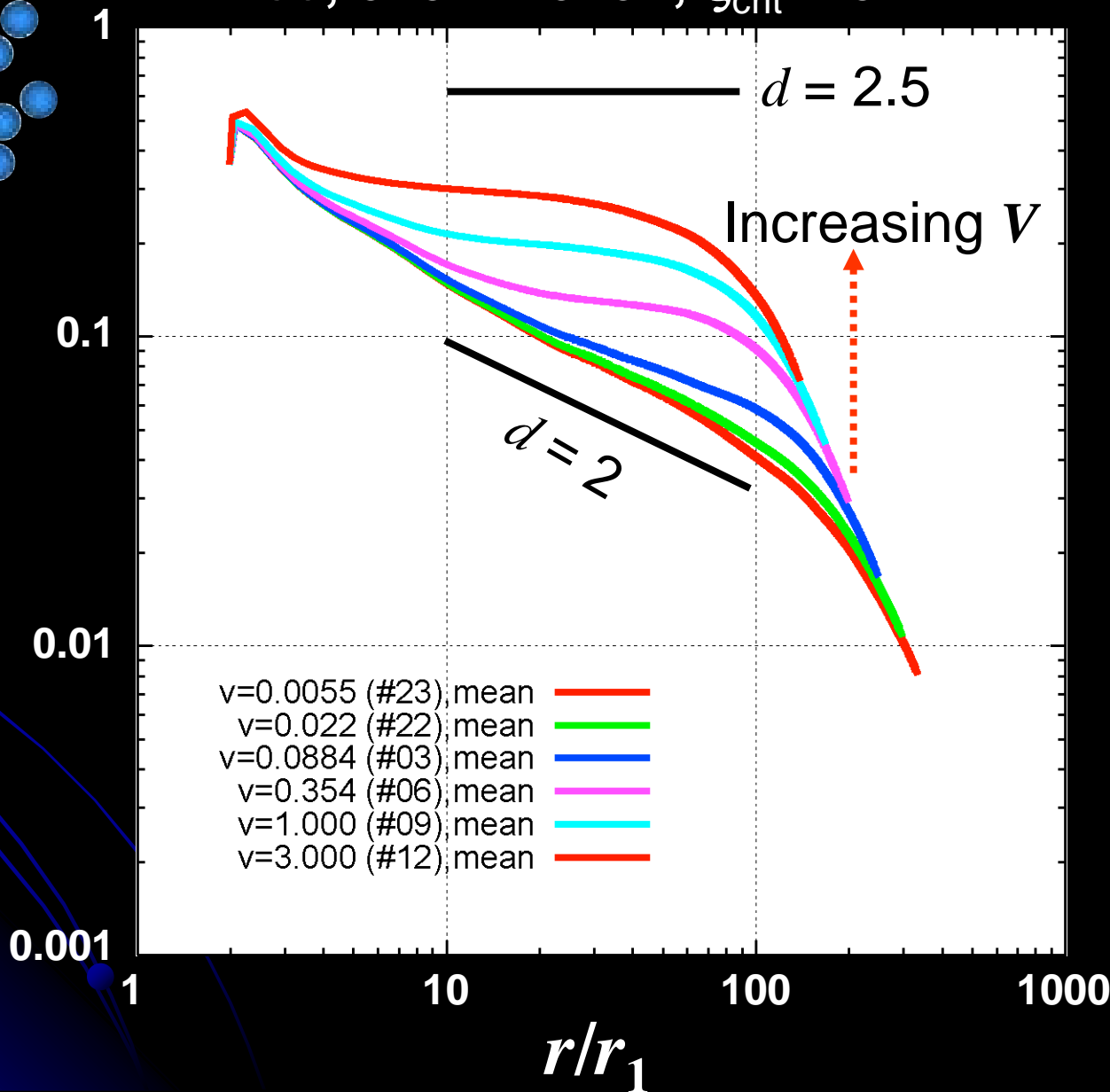
The number of particles $N(<r)$ within r in an aggregate



Ice, 8192 + 8192, $\xi_{\text{crit}} = 8 \text{ \AA}$



$$N / (r/r_1)^{2.5}$$

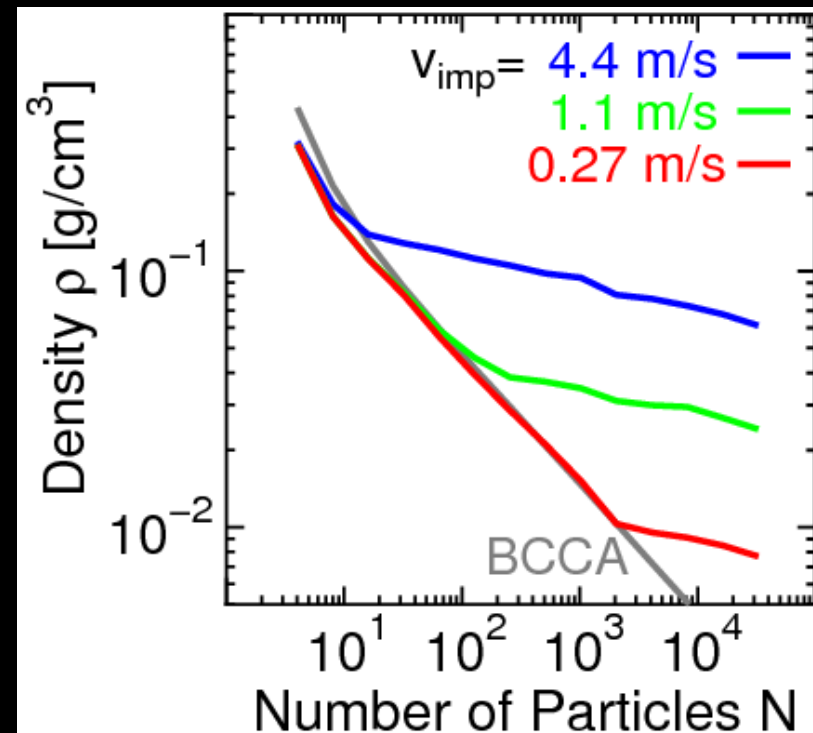


Successive collisions in a BCCA mode



Suyama et al. 2008

- ✓ Fractal dimension ~ 2.5
- ✓ Decrease in density



CG by Dr. T. Takeda, 4D2Uproject, NAOJ

Summary of Compression Process



- 3D BCCA clusters ($d_f \sim 2$) are not fully compressed

- Fractal dimension for max. compression : $d_c \sim 2.5$

- Gyration radius: $r_g \sim r_1 N^{\frac{1}{2.5}} [E_{\text{impact}} / (n_k E_{\text{roll}})]^{-0.1}$

$$\left[r_g \sim r_1 N^{1/d_c} [E_{\text{impact}} / (N E_{\text{roll}})]^{- (1/d_f - 1/d_c)} \right]$$

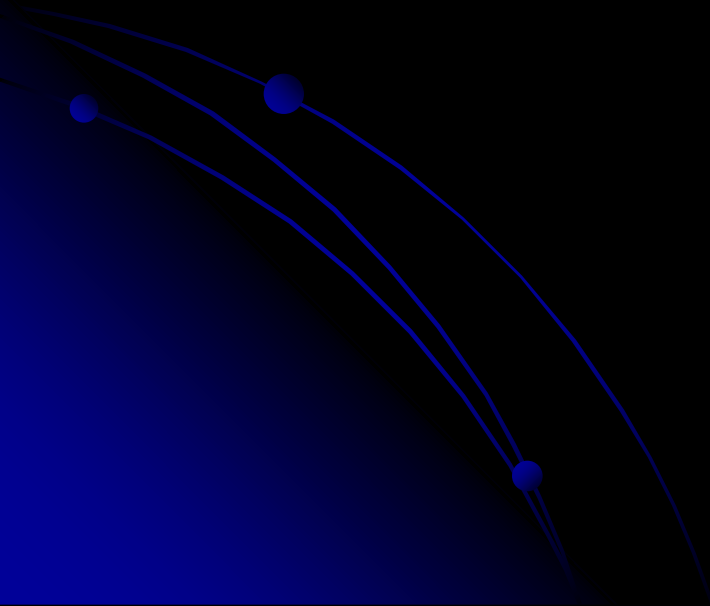
- Successive collisions also lead to $d_c \sim 2.5$

The results for single collisions are applicable.



Collisions between BPCA clusters

: High-velocity collisions



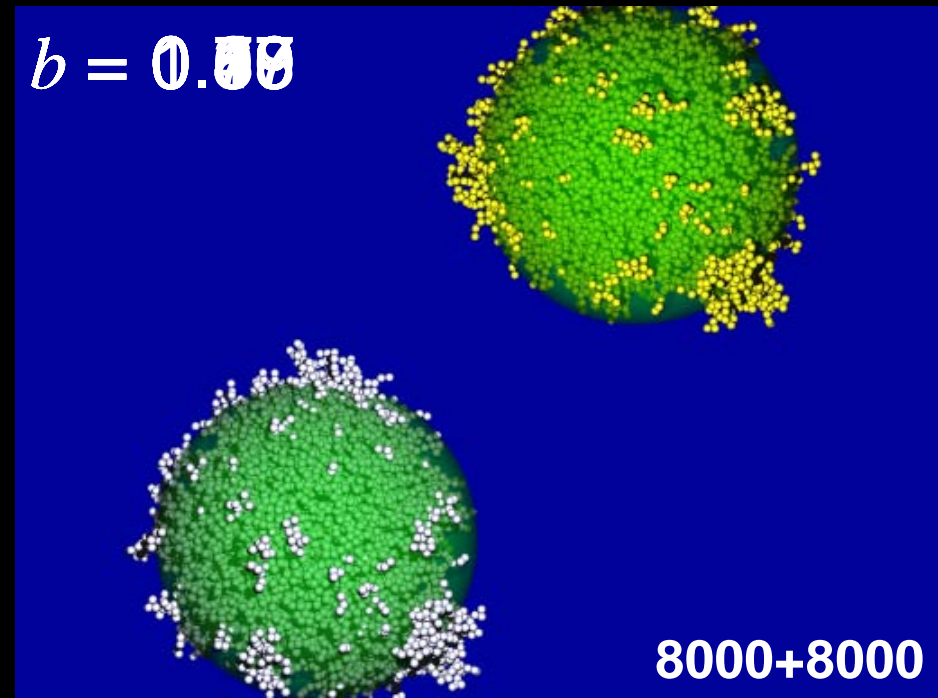
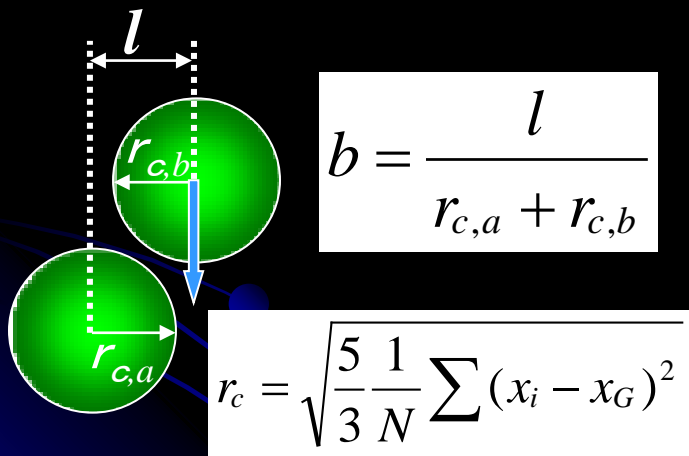
Initial Conditions and Parameters

Collisions of BPCA clusters

✓ BPCA clusters are:

- composed of **500, 2000, or 8000** particles (3 types randomly produced)
- Impact parameter: b (defined by using characteristic radius r_c)

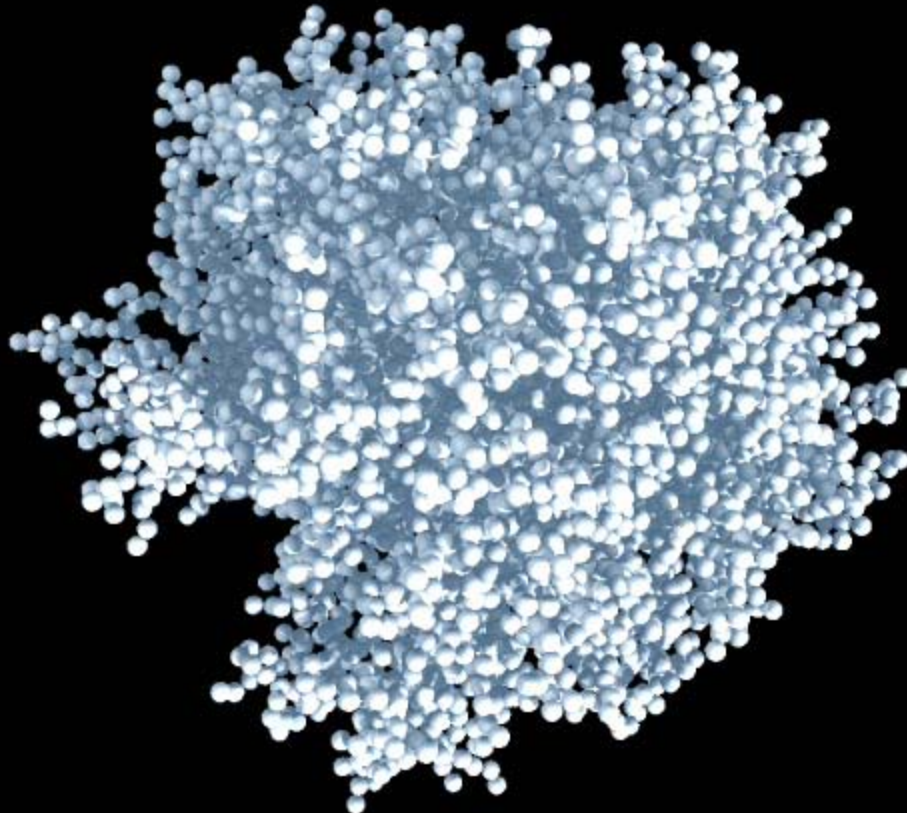
Results are averaged



- ✓ Ice ($E = 7.0 \times 10^{10}$ Pa, $\nu = 0.25$, $\gamma = 100$ mJ/m², $R = 0.1 \mu\text{m}$), critical rolling displace. $\xi_{\text{crit}} = 8 \text{ \AA}$
- ✓ Impact velocity $v_{\text{imp}} = 6 - 260$ m/s



A collision of BPCAs
8000+8000 ice particles ($r=0.1\mu\text{m}$, $\xi_c = 8\text{\AA}$)
Collision velocity = 57 m/s



CG by Dr. T. Takeda,
4D2Uproject, NAOJ

Collisions of BPCA clusters

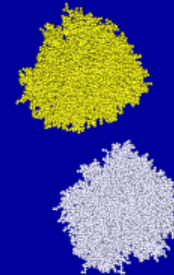


$N=8000+8000$, ice, $\xi_c = 8\text{\AA}$, $v_{\text{imp}} = 70\text{ m/s}$ ($E_{\text{imp}} = 42 NE_{\text{break}}$)

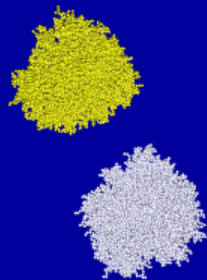
$b = 0$



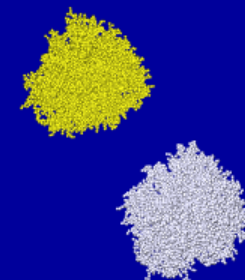
$b = 0.39$



$b = 0.69$

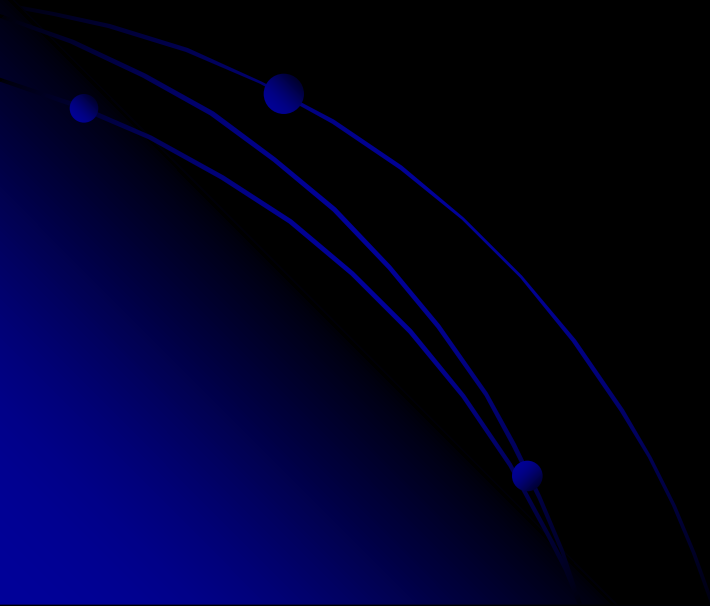


$b = 1.00$

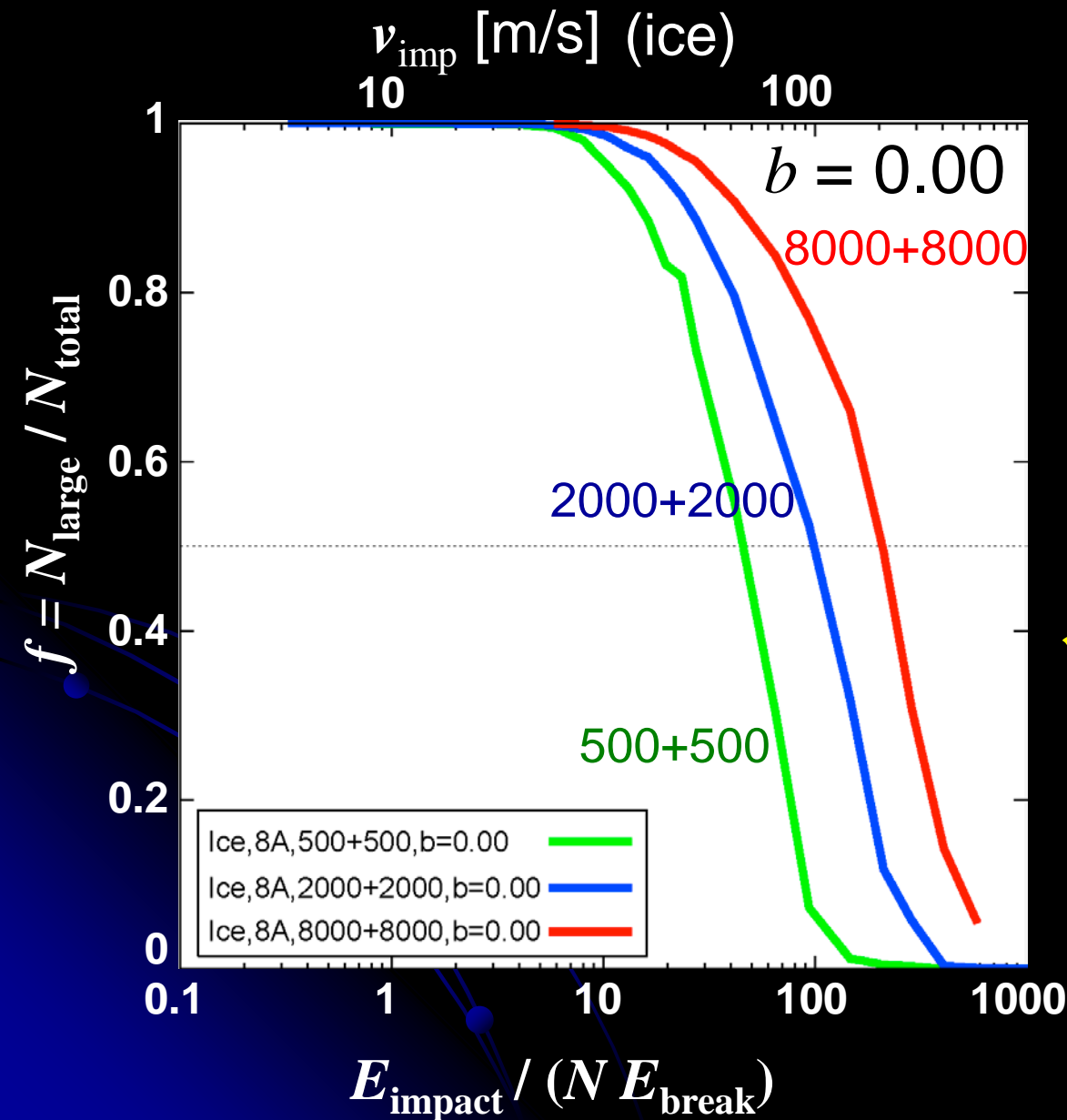




Degree of Disruption: Growth Efficiency



Largest fragment mass N_{large} : *growth efficiency*



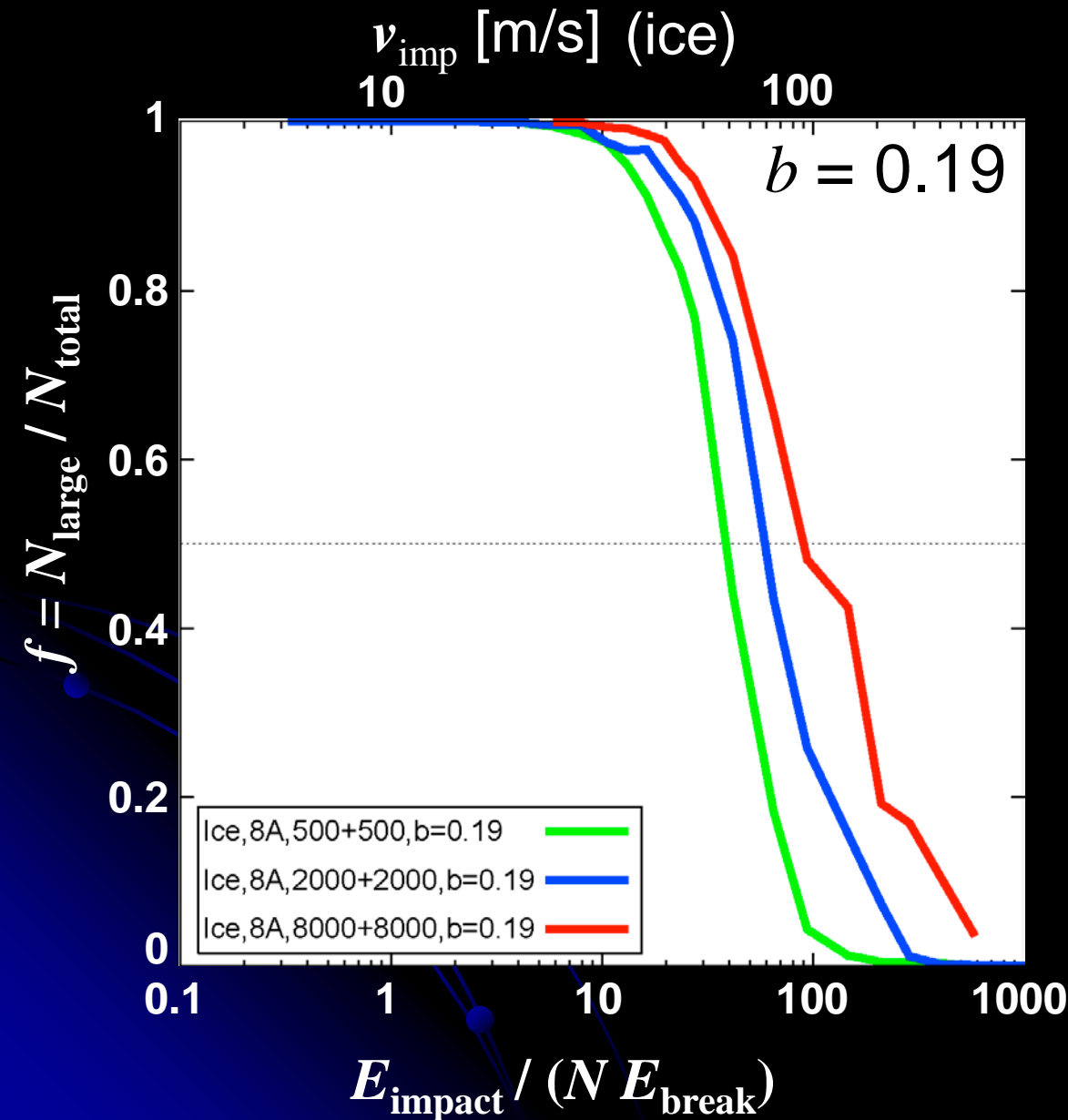
$$f \equiv N_{\text{large}} / N_{\text{total}}$$

: *growth efficiency*

- $f > 0.5 \rightarrow + \text{growth}$
- $f < 0.5 \rightarrow - \text{growth}$

✓ dependent on N

Largest fragment mass N_{large} : *growth efficiency*

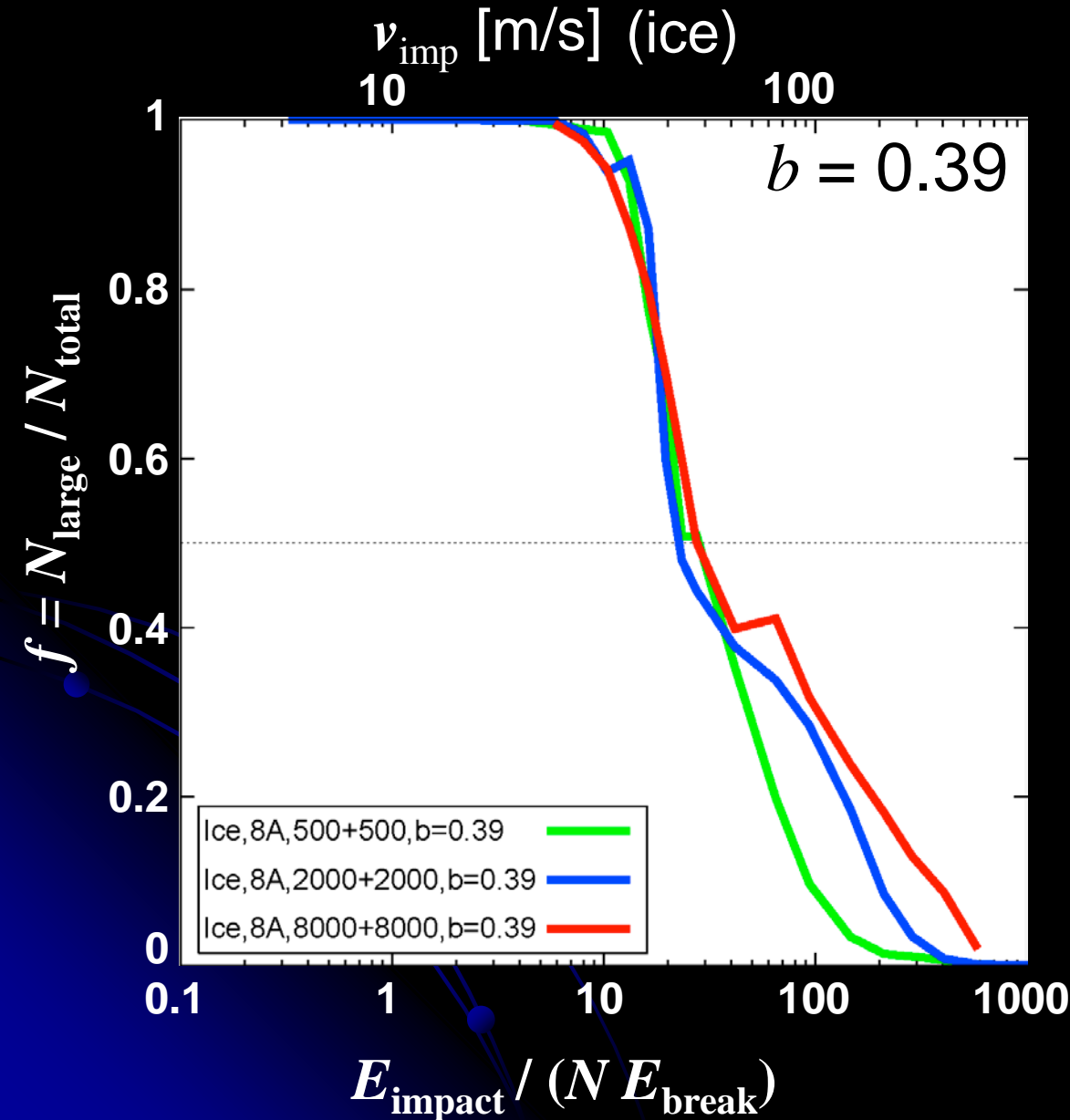


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Largest fragment mass N_{large} : *growth efficiency*

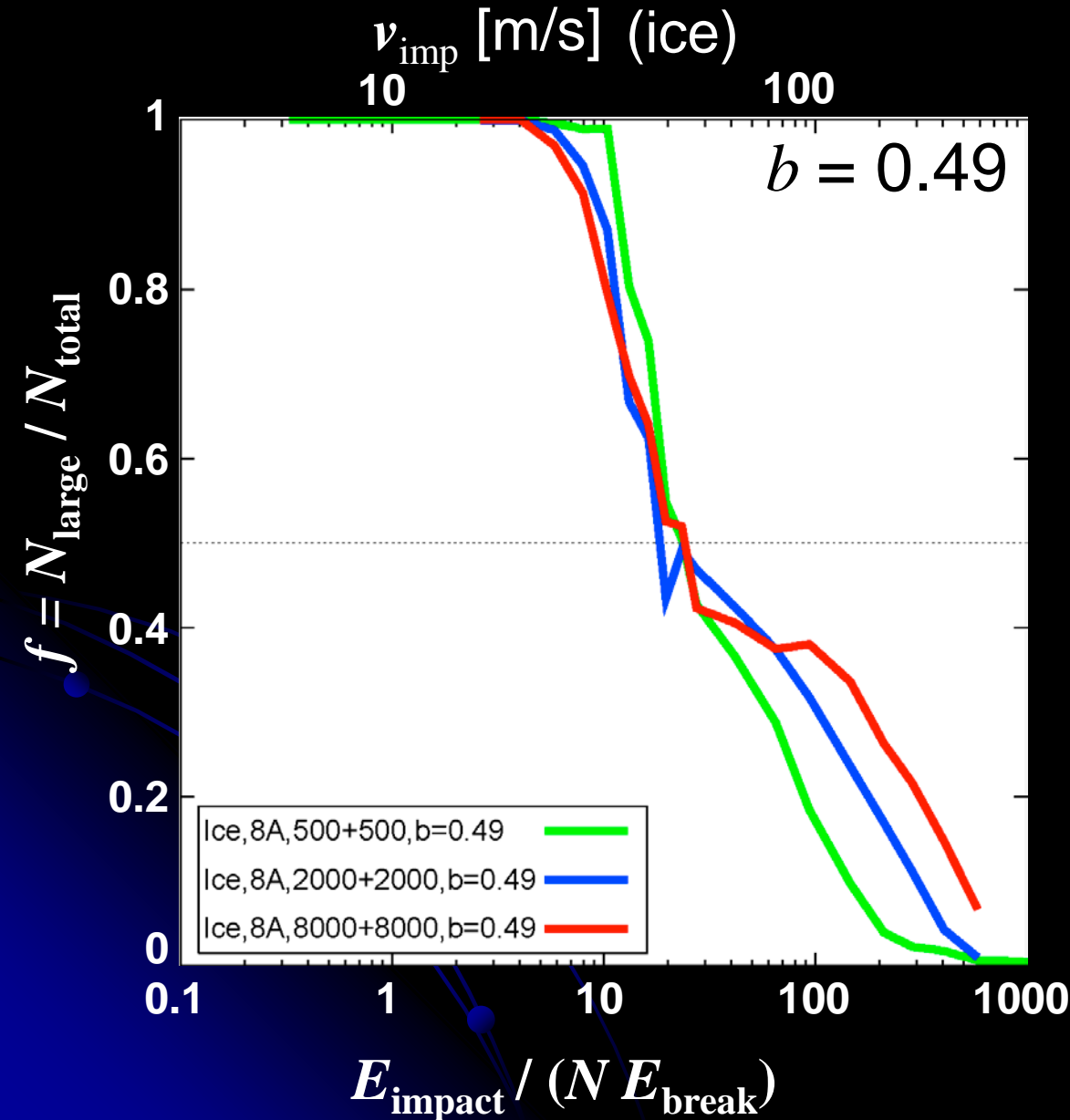


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Largest fragment mass N_{large} : *growth efficiency*

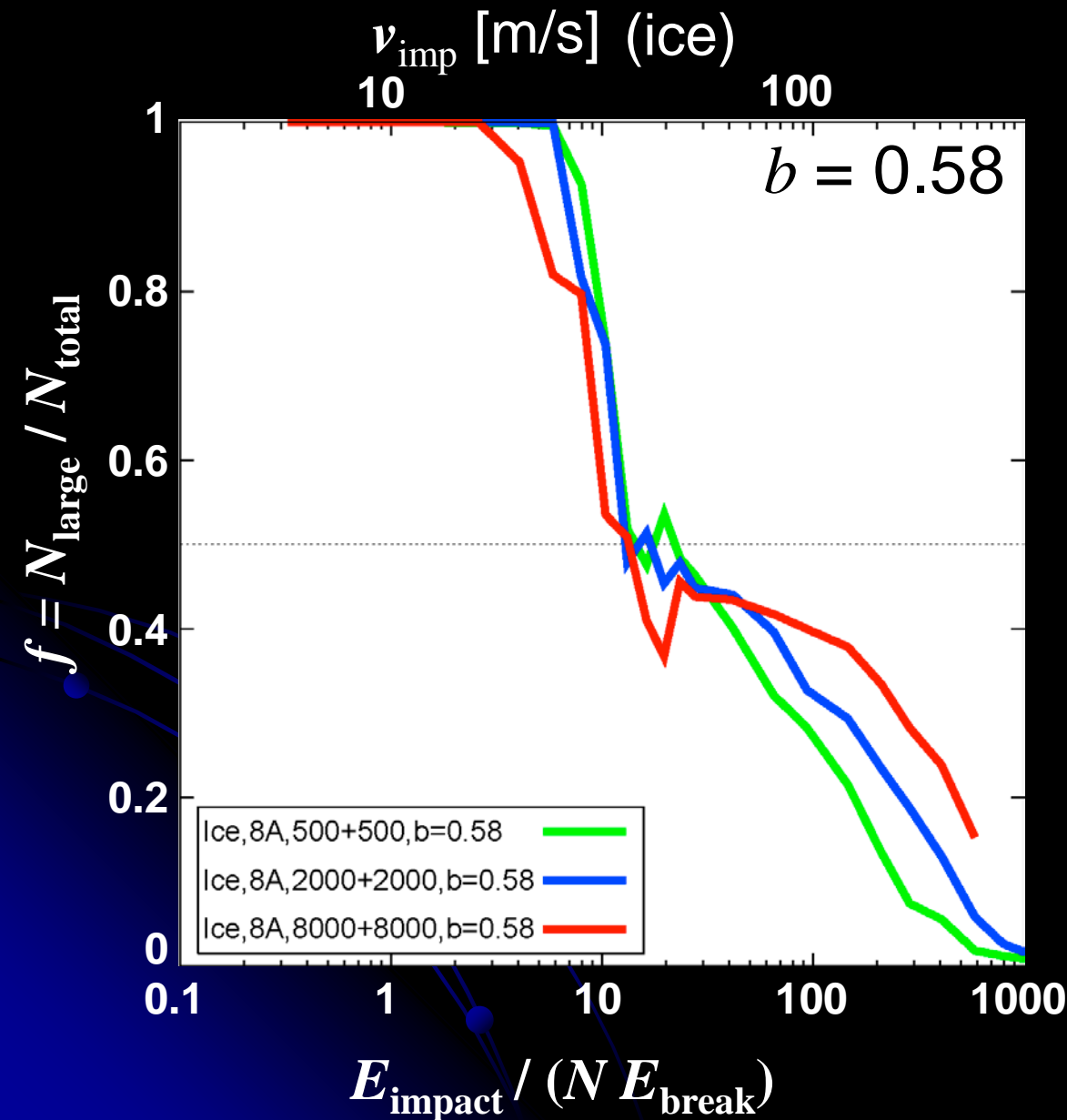


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Largest fragment mass N_{large} : *growth efficiency*

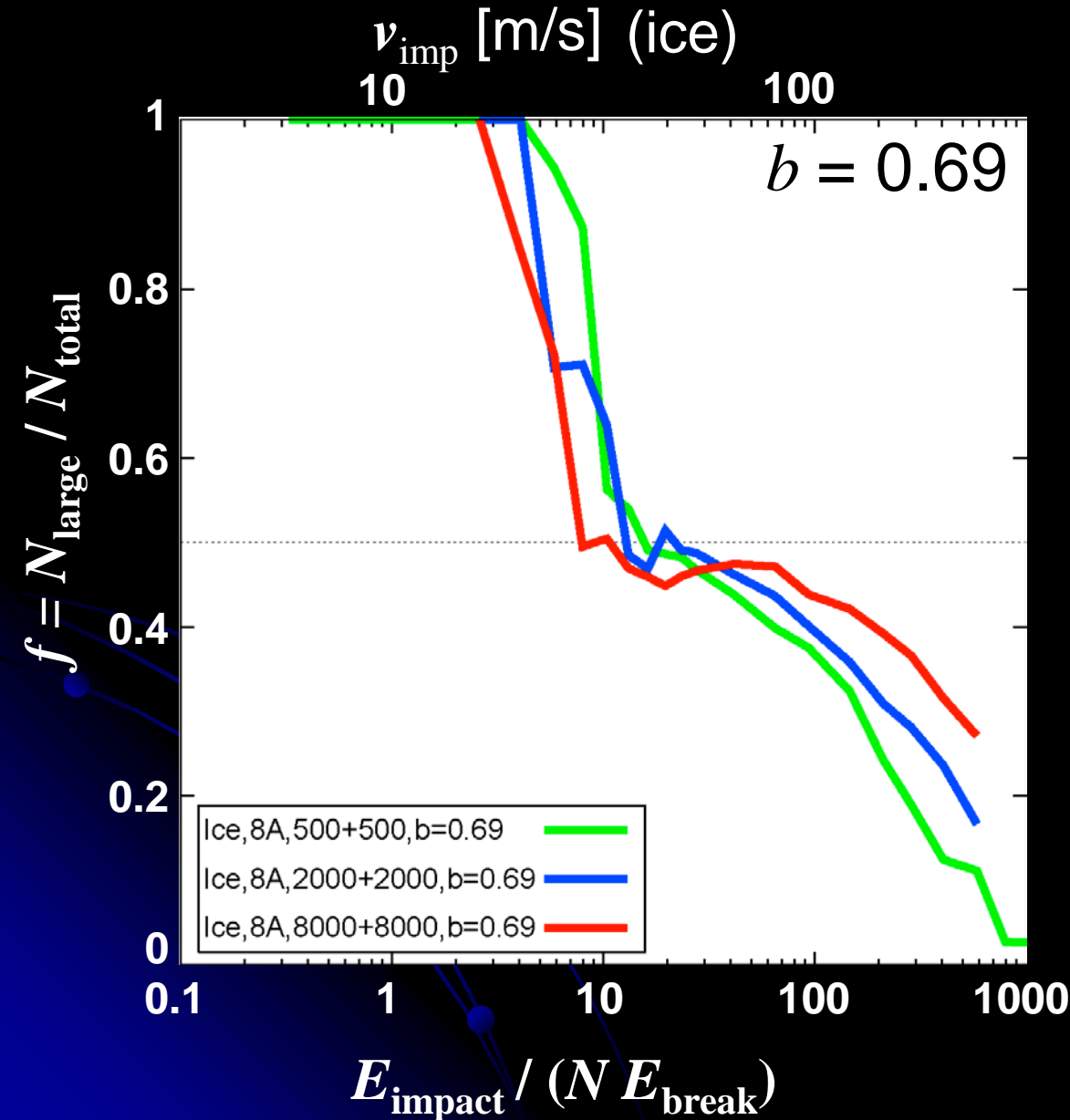


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Largest fragment mass N_{large} : *growth efficiency*

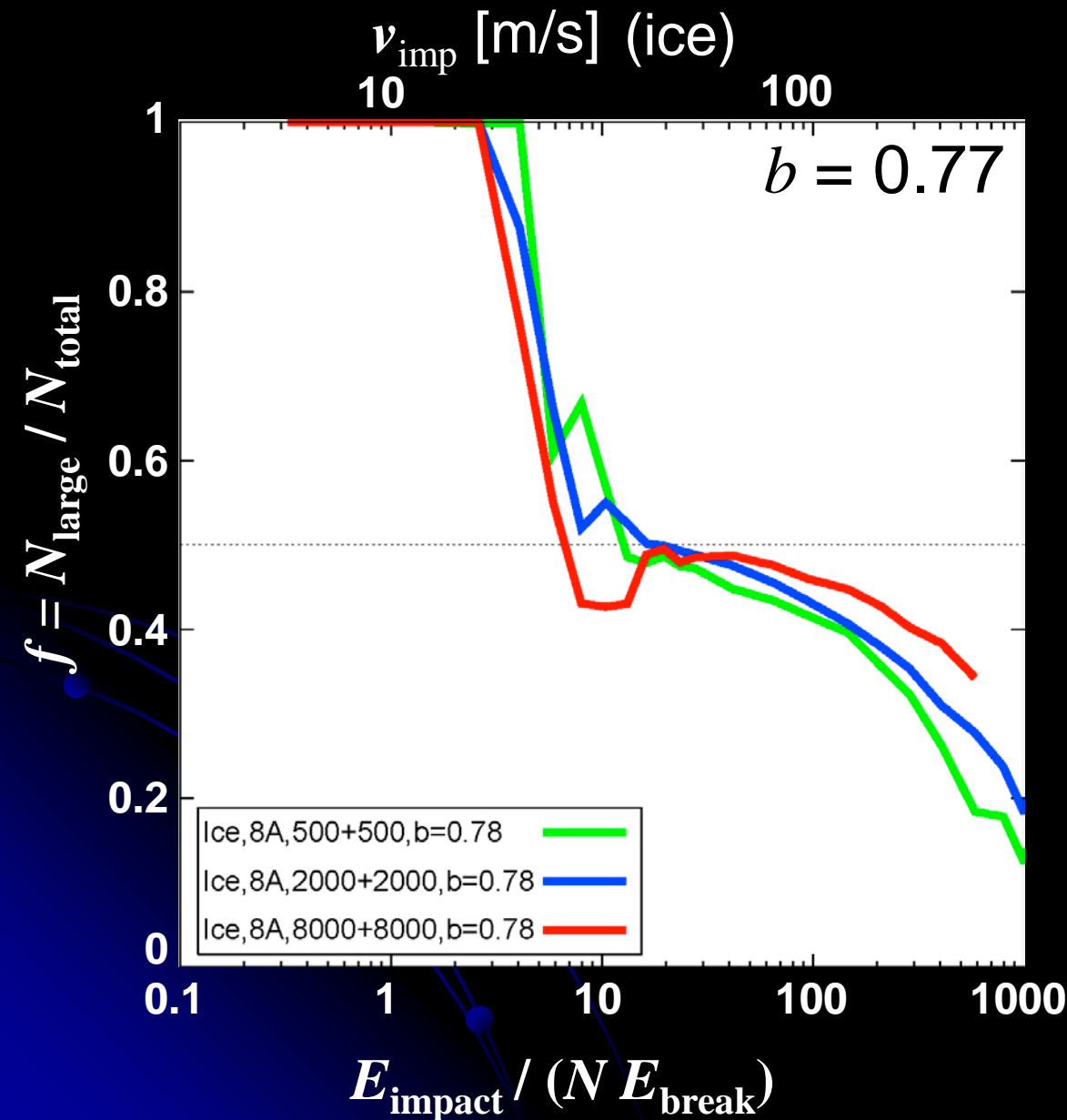


$$f \equiv N_{\text{large}} / N_{\text{total}}$$

: growth efficiency

$f > 0.5 \rightarrow +$ growth
 $f < 0.5 \rightarrow -$ growth

Largest fragment mass N_{large} : *growth efficiency*

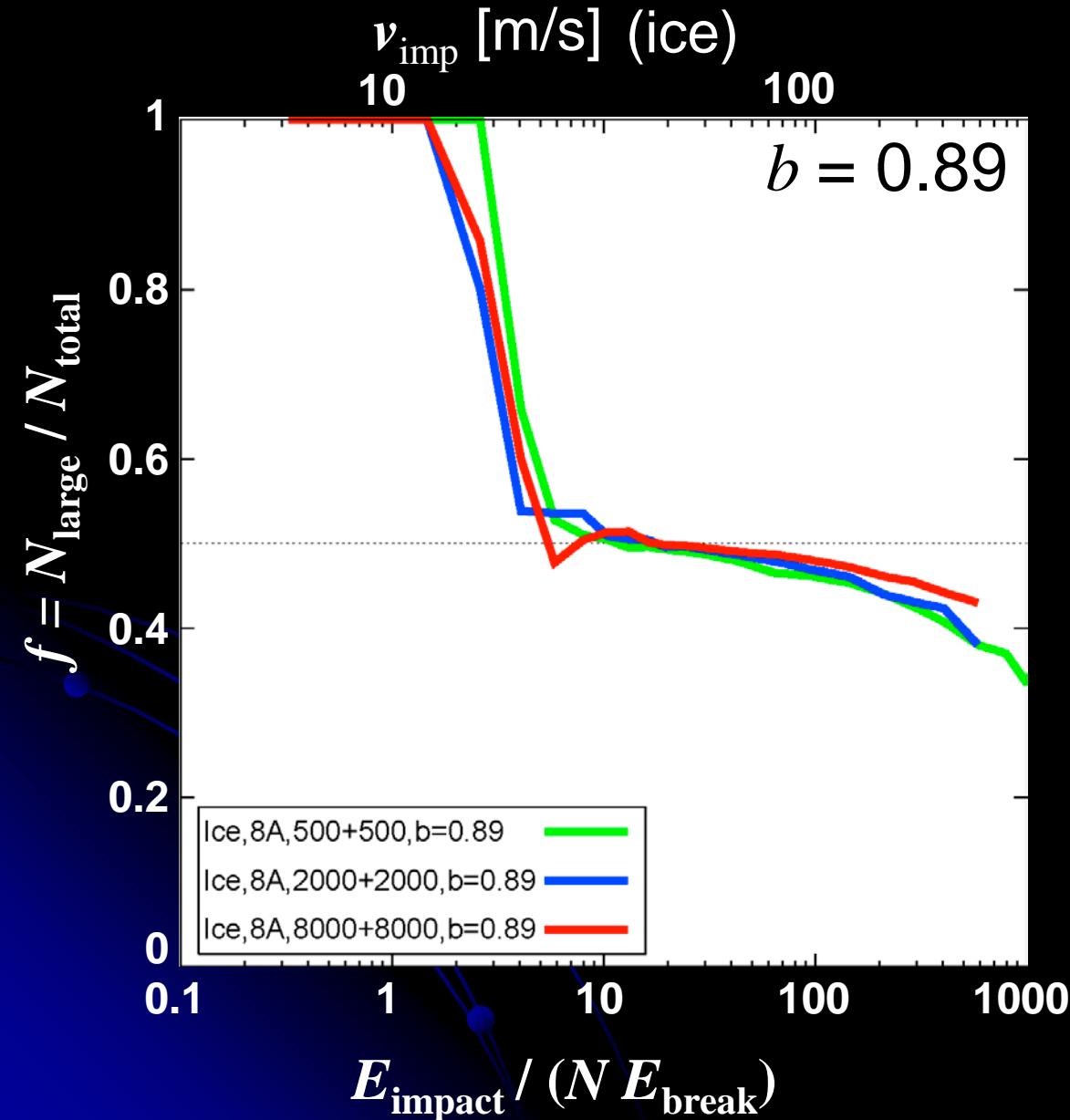


$$f \equiv N_{\text{large}} / N_{\text{total}}$$

: **growth efficiency**

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- $f < 0.5 \rightarrow -$ growth

Largest fragment mass N_{large} : *growth efficiency*

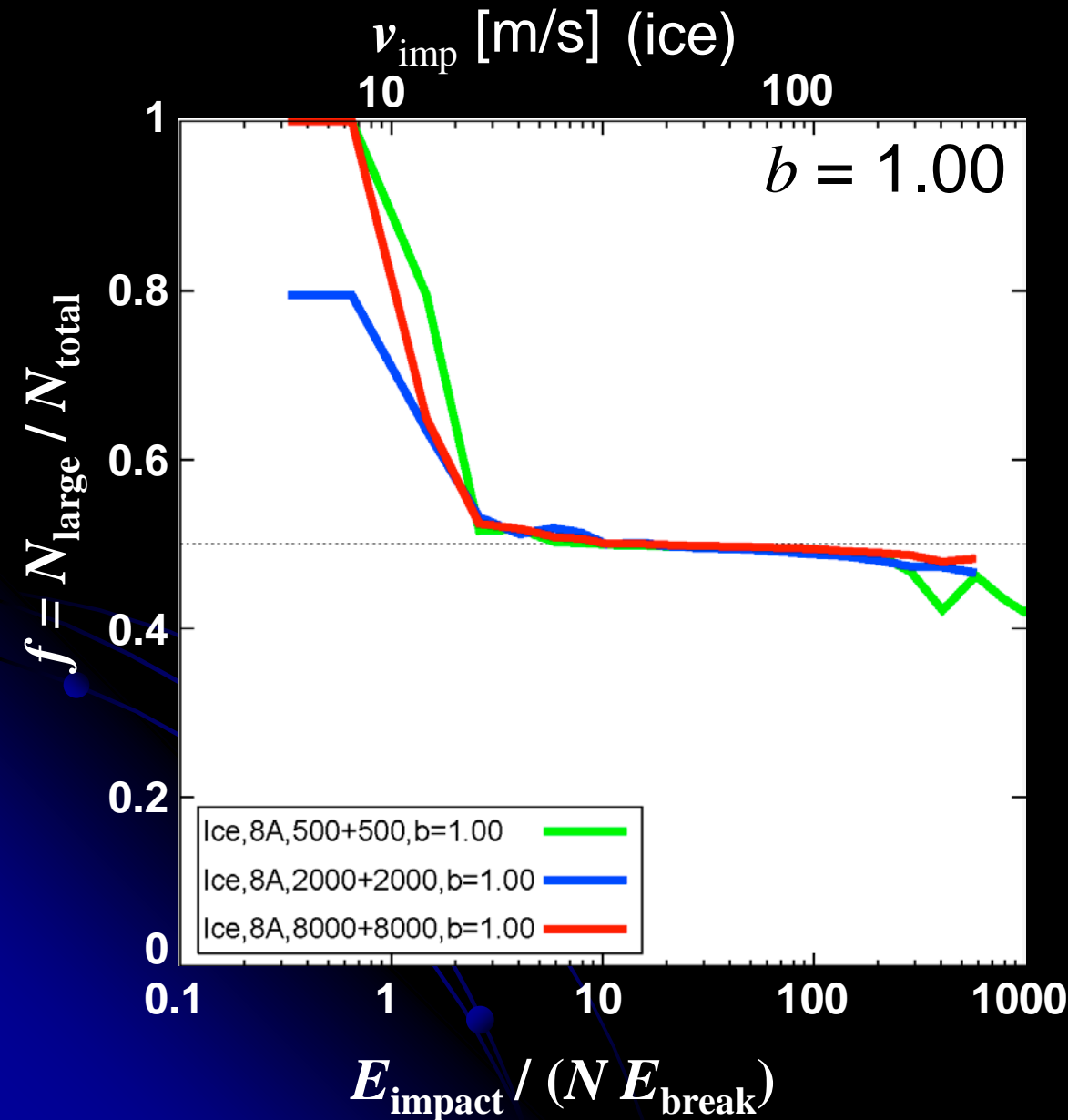


$$f \equiv N_{\text{large}} / N_{\text{total}}$$

: growth efficiency

$f > 0.5 \rightarrow +$ growth
 $f < 0.5 \rightarrow -$ growth

Largest fragment mass N_{large} : *growth efficiency*



$$f \equiv N_{\text{large}} / N_{\text{total}}$$

: **growth efficiency**

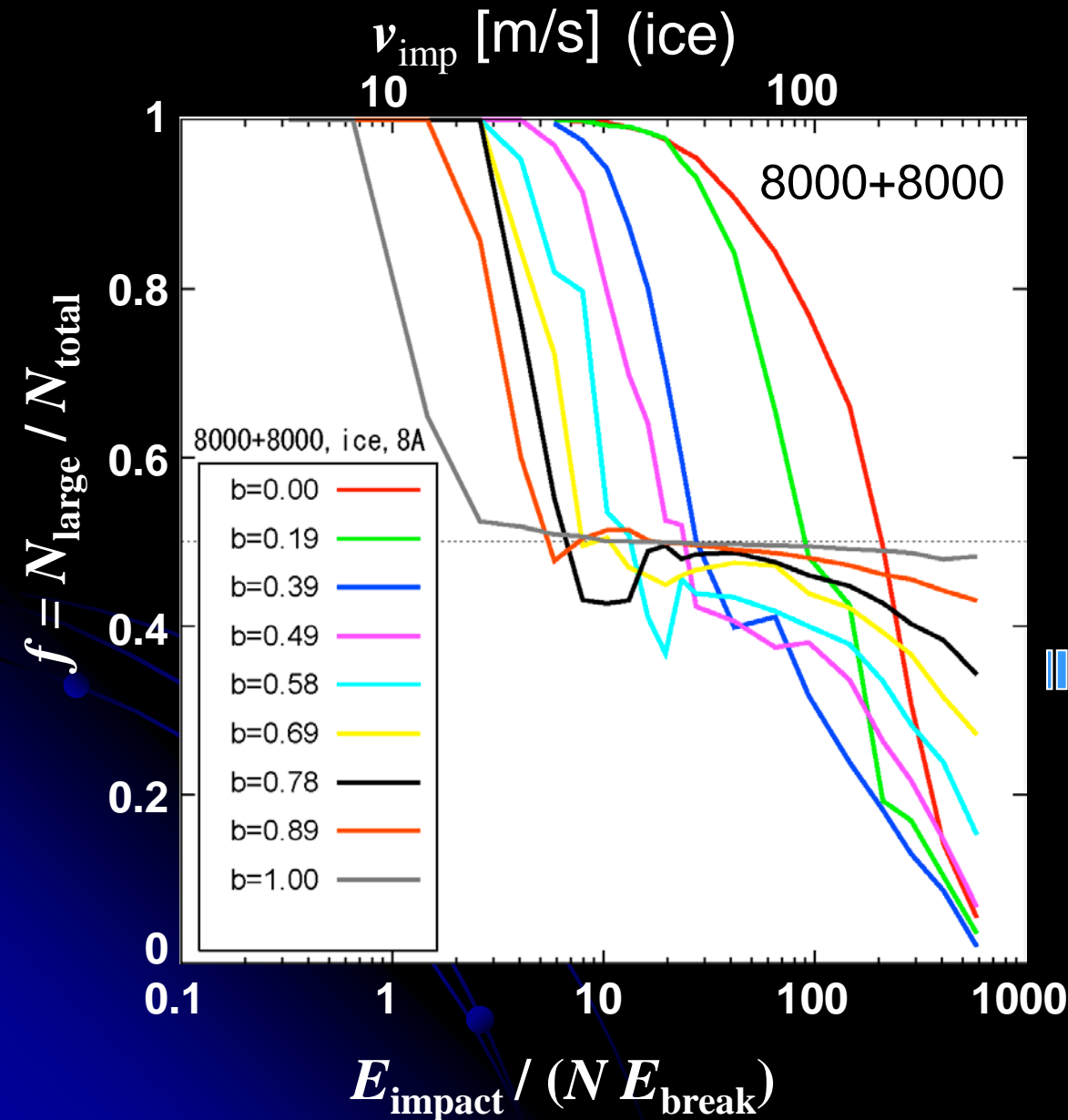
$f > 0.5 \rightarrow +$ growth
 $f < 0.5 \rightarrow -$ growth

✓ Offset collisions



independent of N

Largest fragment mass N_{large} : *growth efficiency*



$$f \equiv N_{\text{large}} / N_{\text{total}}$$

: *growth efficiency*

$f > 0.5 \rightarrow + \text{ growth}$
 $f < 0.5 \rightarrow - \text{ growth}$



Average
weighted by b^2

Growth efficiency averaged

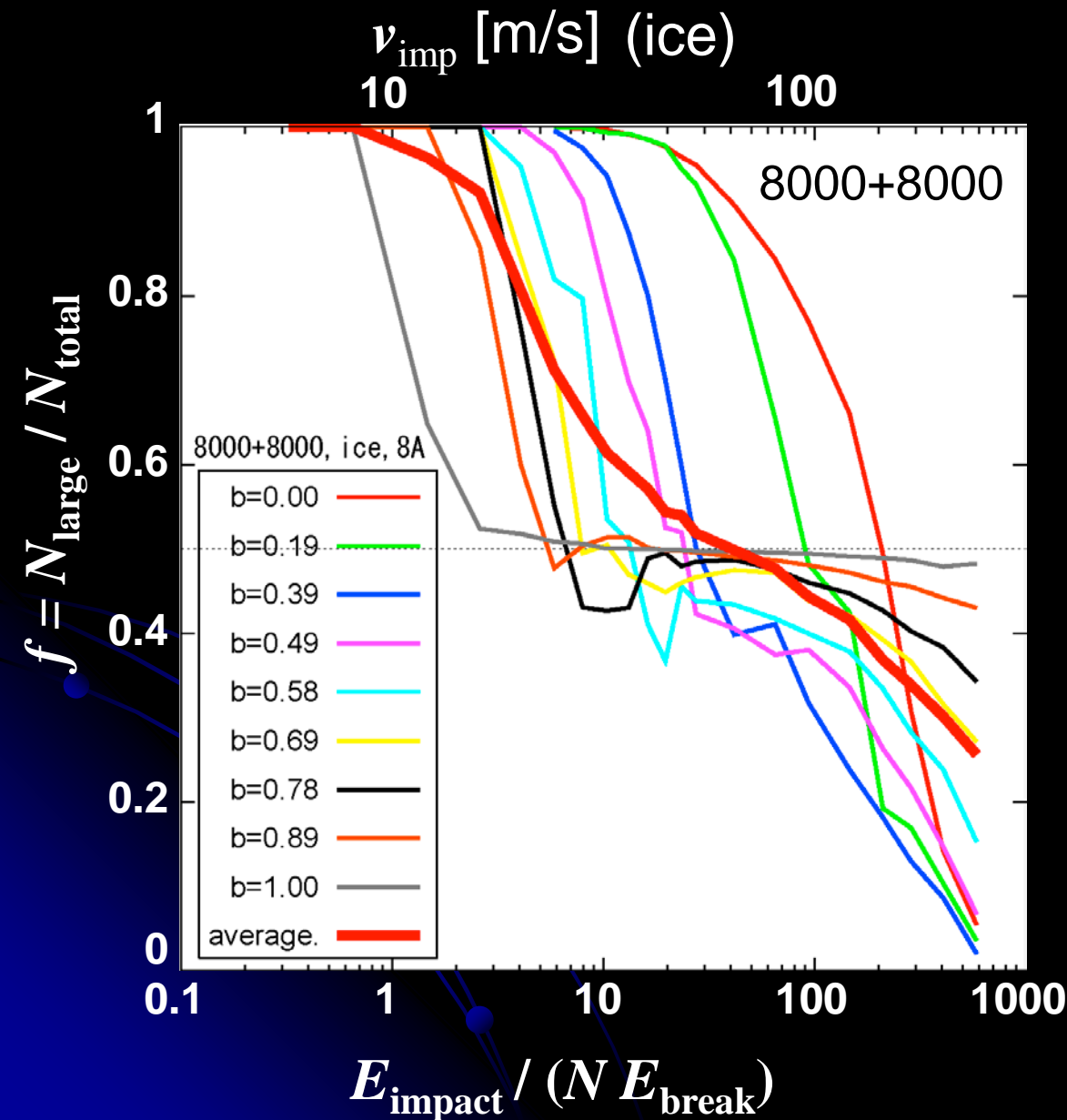
Averaged for b^2

$$f \equiv N_{\text{large}} / N_{\text{total}}$$

: growth efficiency

$f > 0.5 \rightarrow +$ growth

$f < 0.5 \rightarrow -$ growth



Growth efficiency averaged

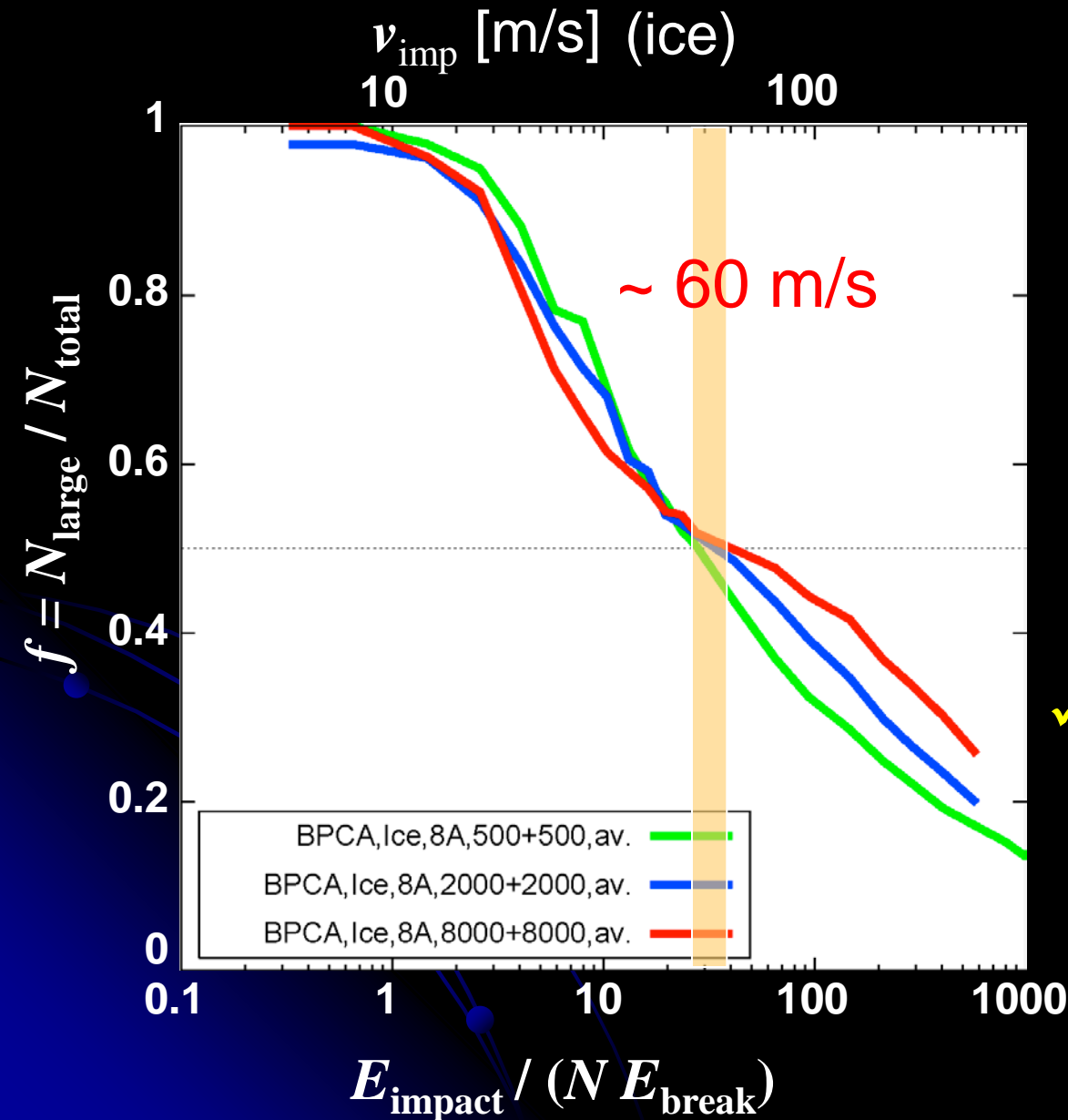
Averaged for b^2

$$f \equiv N_{\text{large}} / N_{\text{total}}$$

f : growth efficiency

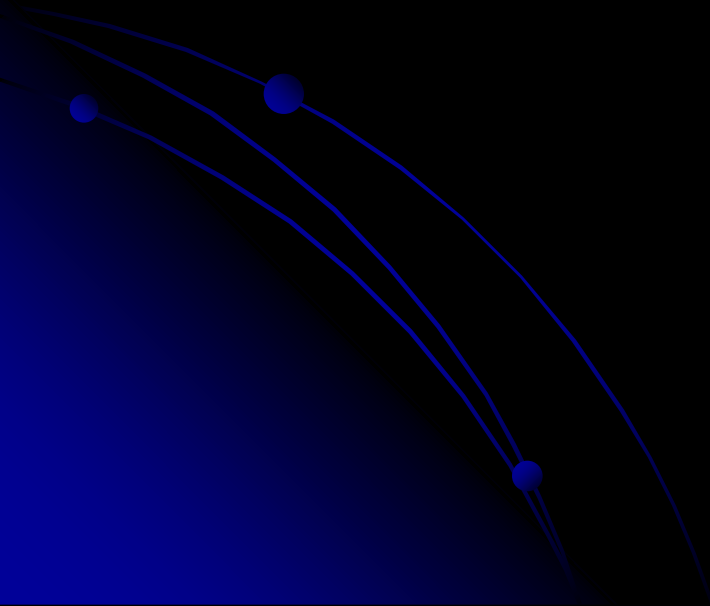
$f > 0.5 \rightarrow +$ growth
 $f < 0.5 \rightarrow -$ growth

✓ small dependence on N





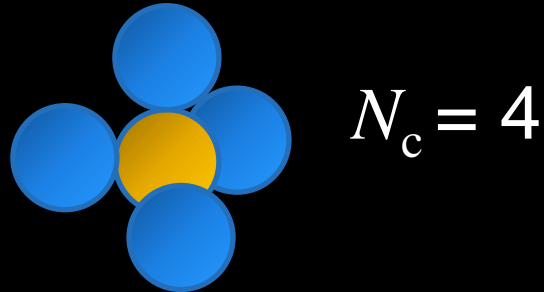
Degree of compression:
Coordination number



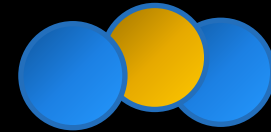
Coordination number N_c

Number of particles in contact with a particle

e.g.,



$N_c = 2$ for *BCCA* and *BPCA*



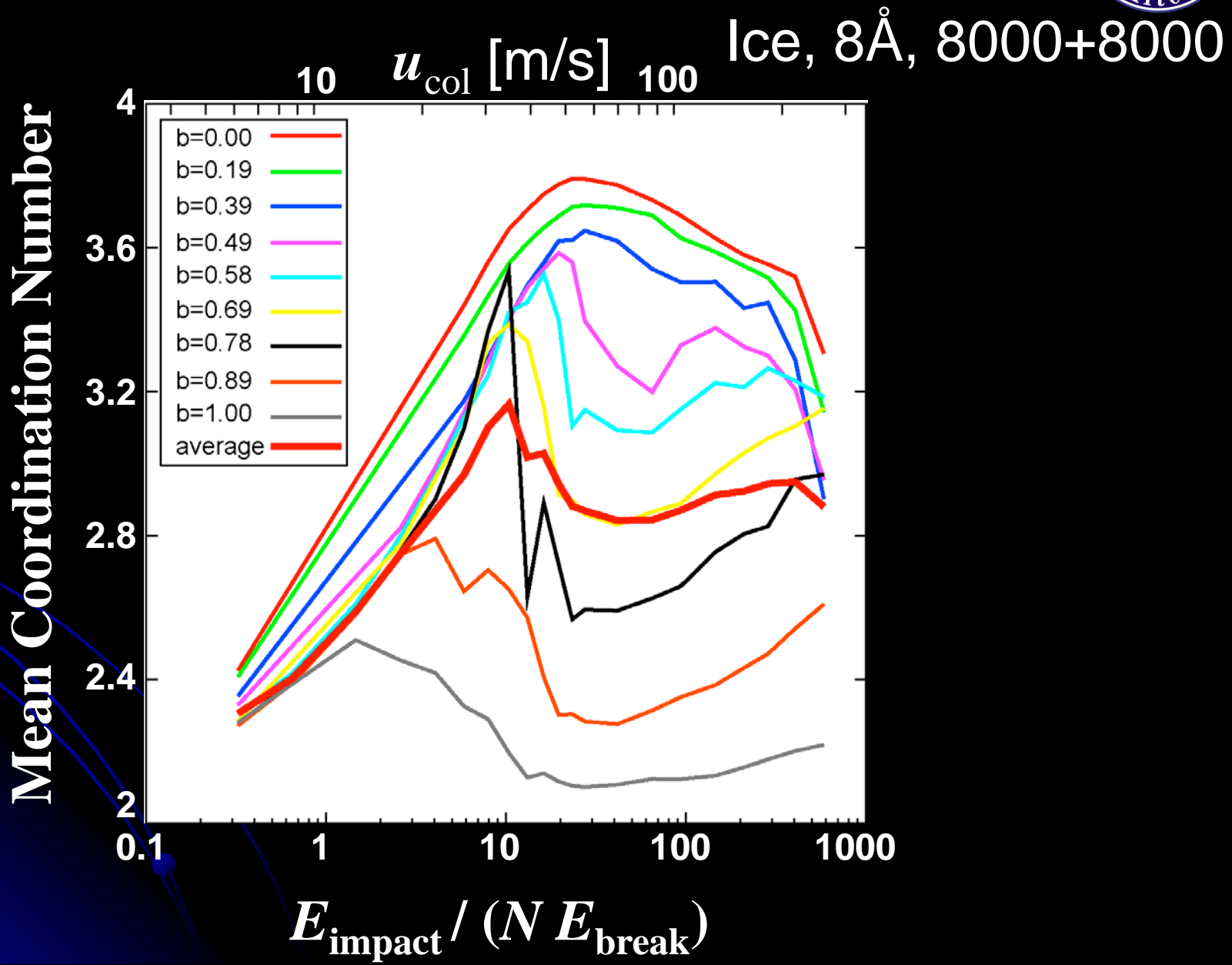
Max. $N_c = 12$ for *close-packing*

An index of compression:

The more compact are aggregates, the larger N_c is.

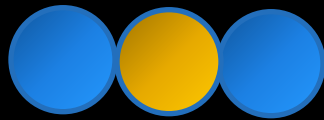
What value of N_c is achieved at *BPCA* collisions?

Coordination number N_c @ BPCA collisions

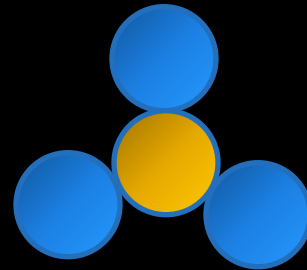


Why $N_c = 4$?

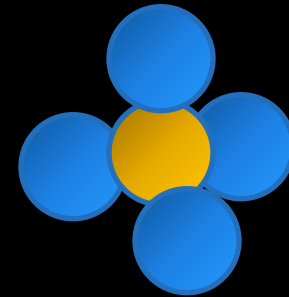
Particles are stable enough with $N_c = 4$ in 3D:



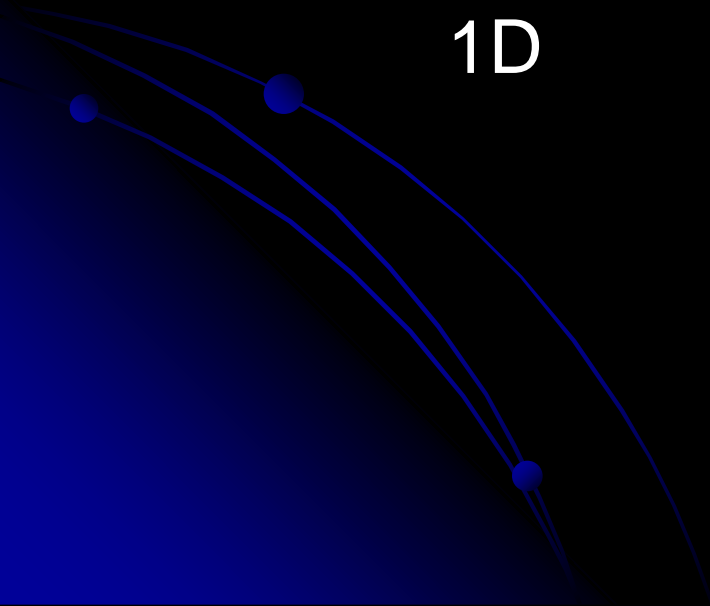
1D



2D



3D





Summary

- Dust aggregates remain fluffly only through collisions.

Fractal dimension ~ 2.5

Coordination number < 4

Very fluffy planetesimals could be formed !?

$\sim 10^{-4}$ g/cc (Suyama et al. 2008)

Other compression processes are required.

- Icy aggregates can grow at collision velocity ~ 50 m/s.

Planetesimals can be formed through collisions of dust.